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## Five-Color Photometry of Bright Stars

\*BRAULIO IRIARTE, HAROLD L. JOHNSON, RICHARD I. MITCHELL, and WIESLAW K. WISNIEWSKI

**A**STRONOMERS have long agreed on the need for more accurate apparent magnitudes of the naked-eye stars. These are, in general, the stars for which other kinds of information are most complete. But as far as brightnesses are concerned, until fairly recently the best available data were still the visual measurements made some 70 years ago at Harvard and Potsdam observatories.

Excellent as this old work was by the standards of its time, the Harvard and Potsdam photometric catalogues of stars failed to meet most later needs. At best, the probable error of an individual observation was nearly  $\pm 0.1$  magnitude, and in the Harvard work there were large systematic errors depending on spectral type, amounting to several tenths of a magnitude for faint red stars. Furthermore, astronomers felt a growing need for measurements of star colors as well as brightnesses.

By the 1940's, improvements in photoelectric photometry provided much more powerful methods. With care and a good sky, stars could be measured to an accuracy of about  $\pm 0.01$  magnitude. By inserting standardized color filters in the photometer, it became possible to measure ultraviolet, blue, and yellow magnitudes of stars. In particular, the filter types used by H. L. Johnson, W. W. Morgan, and D. H. Harris, III, became

\* Braulio Iriarte is at the National Astrophysical Observatory, Tonantzintla, Mexico. Harold L. Johnson and Richard I. Mitchell are at the Lunar and Planetary Laboratory of the University of Arizona, Tucson. Wieslaw K. Wisniewski is also there, on leave of absence from Cracow Observatory in Poland.

### EDITORIAL COMMENT

Eight years ago we published a list of the 50 brightest stars, compiled by Harold L. Johnson from photoelectric measurements (see August, 1957, page 470). Now Dr. Johnson and his colleagues present (on pages 25-31) a greatly enlarged new listing of magnitudes and colors for 1,325 stars, including those easily visible to the naked eye north of declination  $-50^\circ$ .

The new V values are essentially visual magnitudes, but of far greater precision than in any extended lists available up to a few years ago. Good photographic (blue) magnitudes for bright stars can be obtained by adding the V values to the B-V color indexes in the catalogue. Other color data there range from ultraviolet to infrared.

adopted generally, and served to define the new U, B, and V magnitude system.

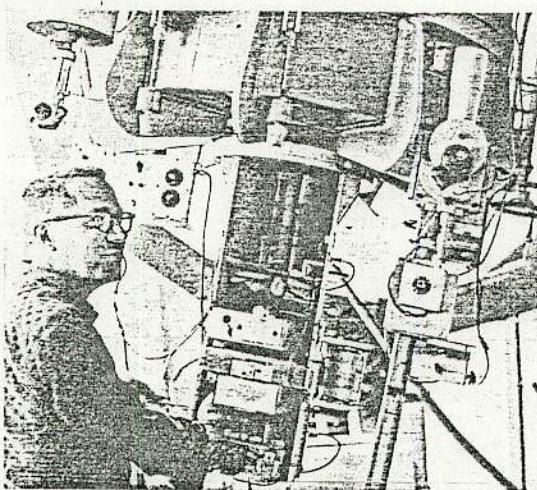
The U magnitudes refer to ultraviolet starlight of around 3600 angstroms wavelength. The blue magnitudes, B, have an effective wavelength of 4300, and may be considered as "photographic" magnitudes. They contain no ultraviolet light short of 3800 angstroms, approximately, in order to minimize the difficulties in comparing and combining the results of different observers. Finally, the V magnitudes refer to yellow light of about 5400 angstroms, and are in effect visual

magnitudes. They average about 0.1 magnitude brighter than the Harvard visual values.

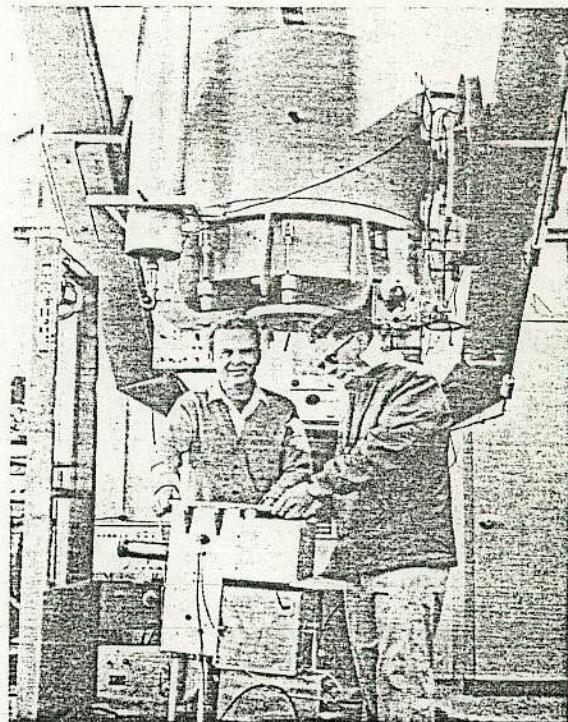
More recently, photoelectric measurements of stars have also been made in the R (red) and the I (infrared) systems, corresponding to around 7000 and 9000 angstroms, respectively. At still longer effective wavelengths are other kinds of magnitudes, determined with lead sulfide cells: J (12,500 angstroms or 1.25 microns), K (2.2 microns), and L (3.6 microns).

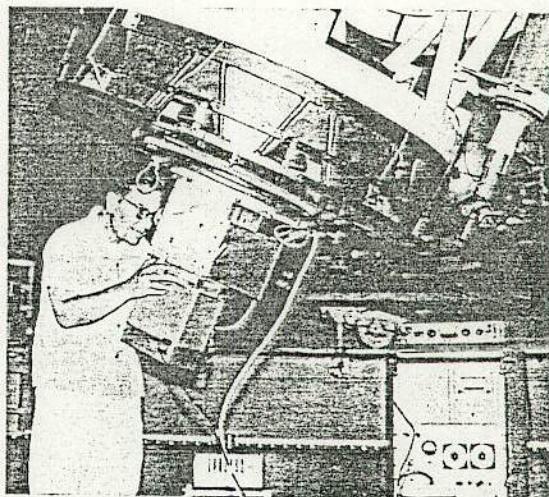
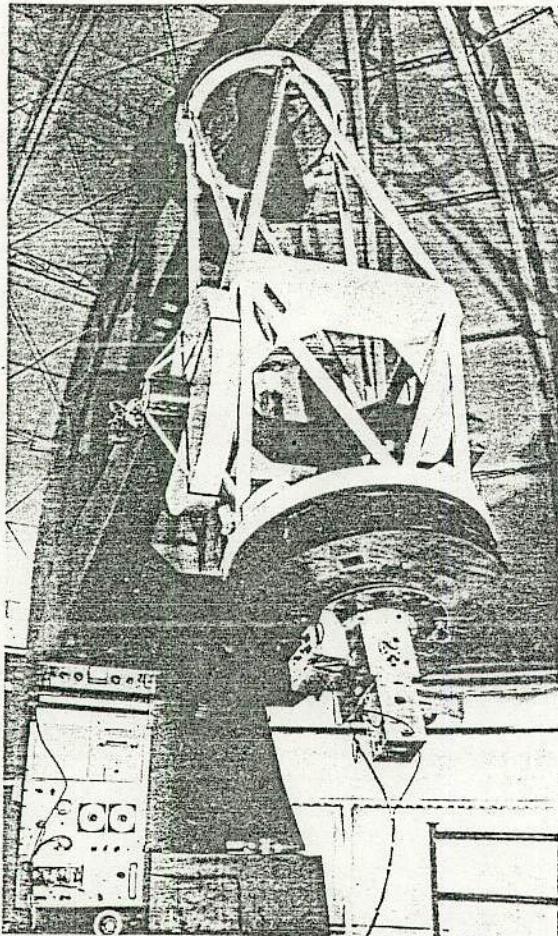
Having so many kinds of accurate stellar magnitudes has greatly widened the concept of star color. Fifty years or more ago, when astronomers were limited to visual and photographic methods, the color index of a star meant merely its photographic magnitude minus its visual magnitude:  $CI = m_p - m_v$ . This index is about 0.0 for white stars (spectral type A0), about +1.0 for orange stars (K0), and +2 or more for very red stars. For many practical purposes today, we can regard the difference B-V as equivalent to the traditional color index.

Similarly, the difference between any two magnitudes for the same star provides another color index. The catalogue that accompanies this article lists U-V, B-V, V-R, and V-I indexes. For the first star in the table, 33 Piscium, these are +1.91, +1.04, +0.77, and +1.32;



Views of the 28-inch Cassegrain reflector at the Catalina station (see *Sky and Telescope* for January, 1964). In David Steinmetz's photograph above, James R. Percy services the bolometer. At right W. K. Wisniewski (left) and Mr. Steinmetz prepare to attach the lead sulfide infrared photometer to the telescope; photograph by Dennis Milon.





Views of the 40-inch Cassegrain reflector at Mexico's National Astrophysical Observatory, located in the village of Tonantzintla, state of Puebla, 80 miles east of Mexico City. Above, Braulio Iriarte places a star in the diaphragm of the photoelectric photometer. At left, the instrument's fork mounting is seen, together with the photometer, and electronic control panel (lower left).

#### COLLECTION OF DATA

The photoelectric observations on which this catalogue is based began in January, 1963, with the 21-inch reflector of the Lunar and Planetary Laboratory in Arizona's Catalina Mountains, 8,250 feet above sea level. That same July, our 28-inch telescope at 8,450 feet came into use on this program. In January, 1965, just two years later, we completed the measurements at this station, in latitude 32° north. At Tonantzintla (7,500 feet elevation, latitude 19° north), observations were begun in February, 1963, and have continued through this May.

Lunar and Planetary Laboratory staff members who took part in the observing included Harold L. Johnson, Richard Mitchell, David Steinmetz, Kent Underwood, Michael Wirick, and Wieslaw E. Wisniewski. The Tonantzintla observations were made by Braulio Iriarte, E. E. Mendoza, and Dr. Johnson.

Descriptions of our UBV and UBVRI photometers have been published in several places, for example in Chapter 7 of *Astronomical Techniques* (edited by W. A. Hiltner, University of Chicago Press, 1962), and *Communications of the Lunar and Planetary Laboratory* (1, 78, 1962). The Catalina photometer contains the original 1P21 photomultiplier tube and filters that defined the UBV system.

The Tonantzintla photometer uses a 7102 photomultiplier. Because of the red sensitivity of this tube, a special filter had to be added to cut the red leaks of the B and V filters. U measurements could not be made.

An important aspect of our program is the automatic data collection system, made possible by a highly routine observ-

adding the first two of them to the V magnitude (4.61) of 33 Piscium gives  $U = 6.52$  and  $B = 5.65$  for this star. Subtracting the last two indexes from V gives  $R = 3.84$  and  $I = 3.29$ . Thus 33 Piscium (a giant star of spectral class K1) appears progressively brighter as we consider its magnitudes from ultraviolet to infrared.

Systematic photoelectric measurement of all these kinds of magnitudes for all naked-eye stars is an enormous task that is still unfinished. The situation until recently was aptly characterized by some remarks in R. O. Redman's address on modern astronomical photometry (*Observatory*, 81, 49-57, 1961).

He said: "My only complaint, as a potential user, is that for quite good astrophysical reasons UBV measurements have been very much concentrated in special regions, e. g. in clusters, and that not enough brighter stars in general have been measured in this system. My hard-hearted colleagues tell me that I must make the measurements myself, but anyone who has tried this type of photometry in the usual English sky will understand why I have a certain reluctance to undertake such work. At present there are more

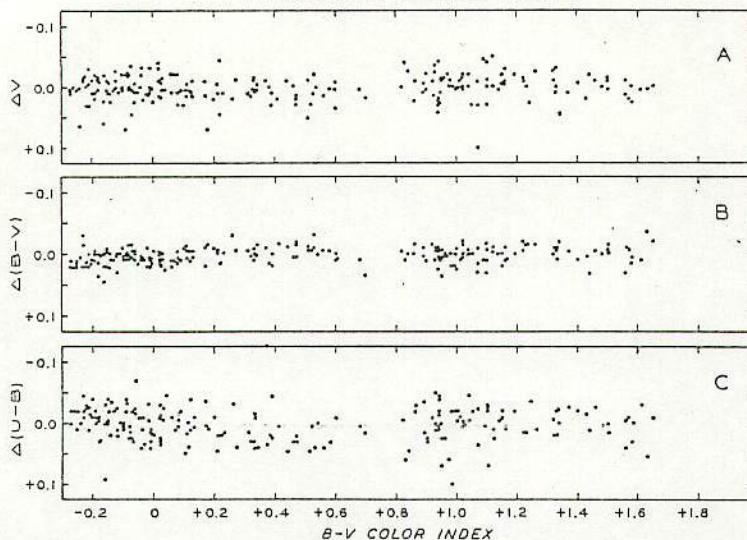
bright stars with first class magnitudes and colours in the southern sky than in the north, thanks to the Cape Observatory [in South Africa]."

We are fortunate that the southwestern United States and much of Mexico have a climate admirably suited for photoelectric photometry, because of the large proportion of extremely clear nights and general freedom from high-level cloudiness.

During the last two years, staff members of the Lunar and Planetary Laboratory (University of Arizona) and astronomers of the Observatorio Astrofísico Nacional (Tonantzintla, Mexico) have participated in a cooperative program of photoelectric photometry of the brightest stars.

Beginning on page 25, we present photoelectric measurements of more than 1,300 bright stars in five broad wavelength bands—U, B, V, R, and I. The stars selected are those brighter than 5.0 on the Harvard visual system. The catalogue is complete to this limit for all of the sky north of declination  $-30^{\circ}$ , and almost complete between  $-30^{\circ}$  and  $-50^{\circ}$ . A few stars fainter than 5th magnitude have also been included.

## CATALINA MINUS CAPE



Comparisons of Catalina measurements and Cape of Good Hope photometry (from Royal Observatory Bulletin 64) for stars between declinations  $-10^{\circ}$  and  $+6^{\circ}$ . In each case the ordinate is the difference Catalina minus Cape.

ing procedure. Beforehand, an IBM card is punched for each star, listing an identification number, the right ascension and declination, spectral type, and visual brightness. (For a faint star, a finder chart may be pasted on the unused portion of the card.) At the time of observation, the identification and coordinates are punched from the card into a paper-tape record. On this tape are then punched the photometric measures of the star, the sky background brightness, and the time. Finally, the observer characterizes the observation as good or bad by means of push buttons.

The tape records from several nights are converted to IBM cards, which are fed into a computer. This computer is programmed to correct the observations for atmospheric extinction and to convert them to the standard UBVRI system. Many astronomers now use computers to reduce their photometric observations.

## SOME FURTHER COMMENTS

The original UBV system is defined by northern stars, many of which cannot be reached by Southern Hemisphere observers. Therefore our new UBV measurements, extending to declination  $-50^{\circ}$ ,

provide additional checks on the photometry in the southern sky.

Recently the Cape Observatory published two lists of UBV magnitudes for most of the bright southern stars (in Bulletin 64 of the Royal Observatory, and in a mimeogram entitled *Photometry from 1961 to 1963*). Both lists are internally consistent and are close to the original UBV system. The accompanying charts comparing Cape and Catalina data show the mutual consistency of these two

series. This comparison is made for equatorial stars that were observed at approximately equal altitudes from Arizona and South Africa.

However, for the bluest stars, the Cape  $B - V$  values (chart B) tend to be more negative than values in the Johnson-Morgan standard system. The Tonantzintla comparison (chart F) confirms this systematic effect for stars from declinations  $-30^{\circ}$  to  $-50^{\circ}$ . The effect is clearest in chart D, in which the Catalina values are compared with those from the Cape mimeograms, this time the  $B - V$  differences being plotted against  $U - B$  values.

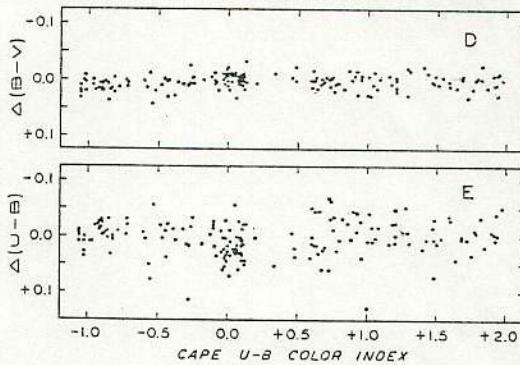
From chart D comes a simple rule for the correction of the Cape  $B - V$  color indexes: If  $U - B$  is negative, apply  $-0.015(U - B)$  to the Cape  $B - V$  value. If  $U - B$  is positive, no correction is applied.

This precept is the only correction needed for the Cape southern photometry. Although the systematic effect seems to come from a difference in the ultraviolet transmissions of the  $B$  (blue) filters, no corrections seem needed for the Cape  $U - B$  measures, since they fit the  $U - B$  system reasonably well.

It should be remembered that the new Catalina-Tonantzintla observations do not define the UBV system. Its formal definition is the data for 108 standard stars, published by Johnson and Harris in the *Astrophysical Journal*, 120, 196, 1954.

According to the new data, Vega (B.S. 7001) is 0.04 magnitude brighter than the standard value. Perhaps the star is variable. Such differences have been carefully checked; they are not due to arith-

## CATALINA MINUS CAPE — 1961-63



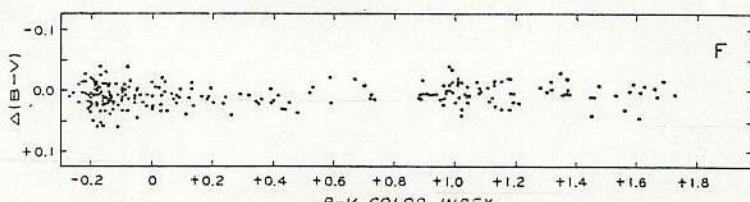
Cape of Good Hope data compared with Catalina measurements as a function of the color  $U - B$ , for stars in the declination zone  $-10^{\circ}$  to  $+6^{\circ}$ . The ordinate in each case is the difference Catalina minus Cape.

metic errors, and instead may represent statistical uncertainty, with some contribution from stellar variability.

The precision of measurement can be evaluated in two ways. The repeatability of the measures of nonstandard stars gives the internal probable error of an observation; comparison with independent data furnishes the external error.

Here are the probable errors of a single observation of each type at Catalina, in units of  $\pm 0.001$  magnitude:  $V$ , 17;  $U - V$ , 20;  $B - V$ , 9;  $V - R$ , 16;  $V - I$ , 16. The corresponding probable errors for a single

## TONANTZINTLA MINUS CAPE



Comparison of Mexican color indexes and Cape values for stars in the declination zone  $-30^{\circ}$  to  $-50^{\circ}$ , well south of the equator. The ordinate is the difference Tonantzintla minus Cape.

Tonantzintla observation are .35, —, 20, 23, 23. (The dash is used because ultraviolet measurements were not made at Tonantzintla, as pointed out previously.)

One of the authors (Johnson) has studied the absolute calibration of the new observations, and his detailed results will appear in the *Communications of the Lunar and Planetary Laboratory*.

All stars have been observed on three or more nights, unless noted otherwise. The tabulated magnitudes and color indexes are weighted means, the individual observations being weighted in inverse proportion to the air mass traversed by the star's light.

For measures south of declination  $-30^{\circ}$ , some Catalina data have been averaged with Tonantzintla data, and some Tonantzintla observations of stars north of  $-30^{\circ}$  have been averaged with Catalina measures.

(Parentheses around a V magnitude in the table indicate a Cape Observatory value used instead of a Tonantzintla one. Because of time variations in the yellow-light extinction at the Mexican station during this work, the Tonantzintla V magnitudes are affected by systematic effects depending on the right ascensions of the stars, and the weight of such a Tonantzintla observation is about a quarter of a Cape one.

Some 70 references were searched to collect the spectral data. W. P. Bidelman of the University of Michigan helped greatly by reading our preliminary list, contributing a number of his own spectral classifications, and providing additional references.

## The Arizona-Tonantzintla Catalogue

### Magnitudes and Colors of 1,325 Bright Stars

THE stars are listed by their numbers (B.S.) in the Yale University Observatory Catalogue of Bright Stars (1964). These are the same numbers used in the Harvard Revised Photometry, published in 1908 as Vol. 50 of the Annals of Harvard Observatory.

An asterisk following the B.S. number indicates a note at the end of the catalogue; *W* (weak) indicates a single observation; *S*, a standard star; *V*, a variable star whose range exceeds 0.08 magnitude, usually with additional information in a note; *D* (double or multiple), two or more neighboring stars measured as one, if the magnitude difference is less than .5.

The star names are given in the next column. Because this table is a computer printout, only capital Roman letters were available to represent the familiar Greek-letter names (see key list). Thus the second star, Alpha Andromedae, appears as ALF AND. In a case such as B.S. 253, UPS 1 CAS is to be read Upsilon-

ALF	Alpha	$\alpha$	NU	Nu	$\nu$
BET	Beta	$\beta$	XI	Xi	$\xi$
GAM	Gamma	$\gamma$	OMI	Omicron	$\circ$
DEL	Delta	$\delta$	PI	Pi	$\pi$
EPS	Epsilon	$\epsilon$	RHO	Rho	$\rho$
ZET	Zeta	$\zeta$	SIG	Sigma	$\sigma$
ETA	Eta	$\eta$	TAU	Tau	$\tau$
THE	Theta	$\theta$	UPS	Upsilon	$\upsilon$
IOT	Iota	$\iota$	PHI	Phi	$\phi$
KAP	Kappa	$\kappa$	CHI	Chi	$\chi$
LAM	Lambda	$\lambda$	Psi	Psi	$\psi$
MU	Mu	$\mu$	OMG	Omega	$\omega$

lon<sup>1</sup> Cassiopeiae. In this column there are some blanks, for stars that have neither Bayer letters nor Flamsteed numbers.

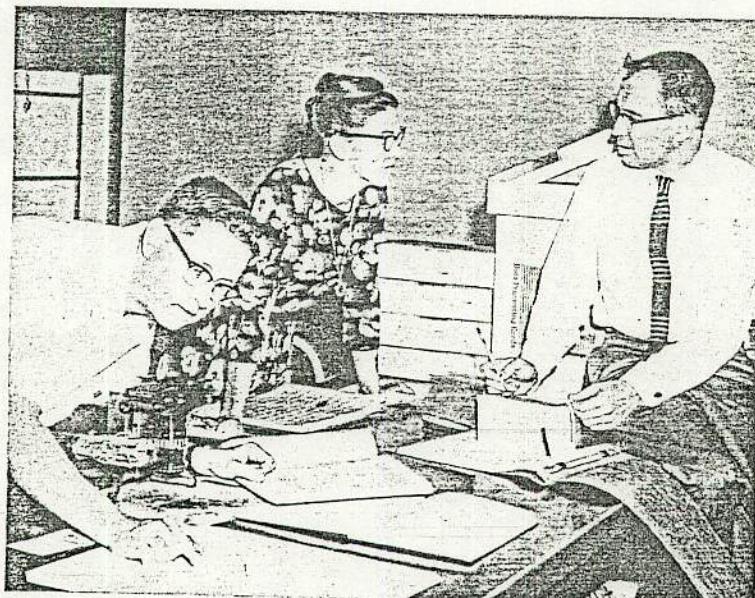
Following the right ascension and declination of each star for epoch 1960 are given the V magnitude and the U — V, B — V, V — R, and V — I indexes, as explained above.

The final column lists spectral types and luminosity classes, whenever possible in the Yerkes (Morgan-Keenan) system. The Roman numeral I indicates a supergiant star, II a bright giant, III a normal giant, IV a subgiant, and V a main-sequence (dwarf) star. When Yerkes classifications are not available, spectral types are taken from the Mount Wilson General Catalogue of Radial Velocities. Suffixes letters in the Mount Wilson types have these meanings: *m*, metallic-line star; *n*, broad lines; *nn*, very broad lines; *p*, peculiar; *e*, with emission lines.

#### B.S. NOTES ON INDIVIDUAL STARS

- 215 Eclipsing var. V spread of our three measures is 0.14 mag.
- 264 Novalike var. Our two measures differ by 0.10 mag. in V.
- 403 Algol-type var. Observed outside of eclipse.
- 424 *Polaris*, Cepheid var. Mean of three observations.
- 493 A single Catalina observation has been averaged with three BVRI ones from Tonantzintla.
- 681 *Mira*. Data in table refer to epoch JD 2,438,675, near time of minimum light. Other observations near maximum, at JD 2,438,473, give: V = 3.42; U — V = 2.52; B — V = 1.68; V — R = 2.15; V — I = 4.02. Visual range about 2.0 to 10.1.
- 1144 Tonantzintla BVRI measures combined with Johnson's published UBV.
- 1151 See note to 1144.
- 1155 Previously suspected var. Our four measures show a V spread of 0.12 mag.
- 1172 See note to 1144.
- 1180 *BU Tauri*, irregular var. V spread of our measures is 0.09 mag. See note to 1144.
- 1239 Algol-type var. Our three measures have a V spread of 0.11 mag.
- 1414 See note to 1144.
- 1496 Previously suspected var. Our three measures show a V spread of 0.09 mag.
- 1547 See note to 1144.
- 1567 Ellipsoidal-binary var. Our three measures show a V spread of 0.07 mag.
- 1612 Algol-type var. Observed outside of eclipse.
- 1781 See note to 1144.
- 1845 *CE Tauri*, semiregular var. Our three measures show a V spread of 0.11 mag.
- 1868 Eclipsing var. Mean of five measures.
- 2061 *Betelgeuse*, semiregular var. Epoch is JD 2,438,317.

(Continued on page 31)



Much of the work on this program is not done at night. At left, David Steinmetz is consulting a catalogue and star maps to prepare finding charts for later use at the telescope. Kathryn Sheffer and Richard Mitchell discuss some observing data received from Tonantzintla, Mexico. Lunar and Planetary Laboratory photograph by Dennis Milon.

## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

B.S.	NAME	R.A.(1960) DEC.	V	U-V	B-V	V-R	V-I	MK SP.	B.S.	NAME	R.A.(1960) DEC.	V	U-V	B-V	V-R	V-I	MK SP.	
J	33 PSC	0 3+3 -56 26	4.61	1.91	1.04	0.77	1.32	K1 III	707 D	IOT CAS	2 25+7 67 13	4.52	0.19	0.13	0.15	0.22	A59	
15 ALF	AND	0 6+3 23 32	2.06	-0.58	-0.11	-0.02	-0.12	BBp	708 RHO	CET	2 24+0 -12 28	4.49	-0.04	-0.02	-0.01	-0.04	B9 V	
21 BET	CAS	0 7+0 58 36	2.28	0.45	0.34	0.31	0.52	F2 IV	715 S	XI 2 CET	2 26+0 8 17	4.23	-0.19	-0.06	0.04	0.00	B9 III	
15 W EPS	PHE	0 7+4 -45 59	(3.87)	0.99	0.75	1.27	K3 III	721 KAP	ERI	2 25+5 -4 53	(4.24)	-0.13	-0.03	-0.16	B5 III			
27 22 AND	AND	0 8+2 45 51	3.04	0.66	0.40	0.41	0.70	F2 II	740 SIG	CET	2 30+2 -15 25	4.74	0.44	0.44	0.41	0.68	F5 IV-V	
39 GAM	PEG	0 11+2 14 58	2.86	-1.08	-0.21	-0.08	-0.31	B2 IV	749 D	OMG	FOR	2 32+1 -28 24	4.89	-0.15	-0.05	0.02	-0.05	B9
45 S CHI	PEG	0 12+5 19 59	4.80	3.50	1.58	1.33	2.45	M2 III	753 S	XI 2 CET	2 33+9 6 42	5.82	1.76	0.97	0.85	1.37	K3 V	
45 V CET	0 12+7 -19 9	4.46	3.61	1.66	1.40	2.54	M1	779 DEL	CET	2 37+4 0 9	4.06	-1.09	-0.21	-0.07	-0.27	B2 IV		
48 THE	AND	0 15+0 38 28	4.61	0.13	0.05	0.08	0.09	A2 V	788 12	PER	2 39+7 40 2	4.92	0.73	0.59	0.50	0.30	F9 V	
48 SIG	AND	0 16+2 34 50	4.50	0.13	0.05	0.08	0.08	A2 V	789		2 38+3 -43 4	(4.75)	-0.08	0.10	0.11	A2 V		
74 IOT	CET	0 17+4 -9 3	3.55	2.39	1.22	0.85	1.44	K2 III	794 IOT	ERI	2 39+1 -40 2	(4.10)	-	1.01	0.79	1.35	KD III	
45 W ALF	PHE	0 24+3 -42 31	(2.83)	-	1.09	0.81	1.40	K0 III	799 THE	PER	2 41+5 6 49	4.13	0.50	0.49	0.46	0.76	F7 V	
102 KAP	PHE	0 24+3 -43 54	(3.22)	-	0.18	0.14	0.22	A7 Vn	801 35	ARI	2 41+1 27 32	4.67	-0.75	-0.13	-0.01	-0.14	B3 V	
125 W ETA	SCL	0 25+9 -13 14	4.54	-	1.62	1.59	2.97	H5 IIID	804 D GAM	CET	2 41+2 3 4	3.47	0.17	0.10	0.11	0.16	A2 V	
125 O LAM	CAS	0 29+5 54 18	4.74	-0.45	-0.10	0.01	-	B8	811 PI	CET	2 42+2 -14 2	4.25	-0.59	-0.14	-0.01	-0.19	BT V	
130 KAP	CAS	0 30+7 62 43	4.16	-0.67	0.13	0.15	0.22	B1 Ia	813 HU	CET	2 42+8 9 57	4.28	0.39	0.31	0.31	0.50	FO IV	
133 ZET	CAS	0 34+7 53 41	3.66	-1.07	-0.18	-0.05	-0.28	B2 V	818 TAU 1	ERI	2 43+2 -18 44	4.46	0.48	0.43	0.70	F6 V		
134 PL	AND	0 34+7 33 30	4.36	-0.71	-0.16	-0.04	-0.16	B5 V	824 39	ARI	2 43+5 29 5	4.52	2.14	1.11	0.80	1.38	K1 III	
133 FPS	AND	0 36+4 29 6	4.39	1.34	0.87	0.68	1.19	G8 IIID	825		2 46+5 56 55	6 26	1.40	0.87	0.82	M5 IA		
125 DEL	AND	0 37+2 39 39	3.30	2.77	1.30	0.92	1.58	K3 III	834 ETA	PER	2 47+8 55 44	3.79	3.59	1.69	1.23	2.12	K3 b	
168 ALF	CAS	0 38+2 56 19	2.22	2.30	1.17	0.79	1.38	K0 I-III	838 41	ARI	2 47+6 27 6	3.63	-0.48	-0.10	-0.00	-0.11	BB V	
179 XI	CAS	0 39+8 50 18	4.81	-0.76	-0.10	-0.00	-0.12	B2 V	840 16	PER	2 48+0 38 9	4.24	0.43	0.35	0.30	0.52	F2 III	
130 MU	PHE	0 39+4 -16 10	(4.58)	-	0.95	0.75	1.27	G8 III	841 BET	FOR	2 47+4 -32 36	(4.45)	-	0.98	0.76	L-130	G6 III	
188 BET	CET	0 41+6 -18 12	2.00	1.88	1.00	0.72	1.24	K1 III	843 17	PER	2 49+0 34 54	4.54	3.50	1.56	2.19	2.18	K3 III	
193 OMI	OMI	0 42+5 48 4	4.50	-0.59	-0.06	0.05	-0.08	B2 V	850 TAU 2	ERI	2 51+2 -21 10	4.77	1.51	0.90	0.70	1.17	KO III	
174 PHI	CET	0 42+2 -10 50	4.75	1.84	1.00	0.74	1.24	K0 III	854 TAU PER	2 51+4 52 36	3.95	1.20	0.75	0.63	1.07	EGL COMP.		
215 V ZET	AND	0 45+2 24 3	4.06	1.99	1.12	0.84	1.44	K1 III	854 ETA	ERI	2 54+5 -9 3	3.87	2.11	1.12	0.79	1.37	K1 III-IV	
219 D ETA	CAS	0 46+5 57 37	3.43	0.62	0.58	0.30	0.80	G0 V	875 S		2 54+6 -3 52	5.17	0.13	0.08	0.11	0.18	A1 V	
224 DEL	PSC	0 46+6 7 22	4.24	3.39	1.91	1.17	2.04	K5 III	879 PI	PER	2 56+2 39 30	4.70	0.18	0.06	0.14	0.25	A2 V	
226 NU	AND	0 47+6 40 52	4.52	-0.52	-0.15	-0.03	-0.18	B5 V	882 24	PER	2 56+6 35 1	4.94	2.53	1.25	0.87	1.53	K2 III	
244 UPS	DEL	0 50+6 60 54	4.83	0.65	0.53	0.48	0.77	FB V	896 LAM	CET	2 57+6 8 45	4.70	-0.57	-0.12	-0.02	-0.16	S8 III	
248 20 CET	0 51+0 -1 21	4.78	3.49	1.57	1.23	2.14	K0 III	897/8 THE	ERI	2 56+8 -6 28	(2.90)	*	0.11	0.15	0.22	A3 V		
253 UPS	1 CAS	0 52+8 58 46	4.84	2.48	1.22	0.92	1.53	K2 III	911 91	ALF	CET	3 0 3.2	3.36	2.32	3.57	1.63	2.51	K2 III
254 V GAM	PHE	0 53+4 60 30	2.41	-1.21	-0.11	-0.09	0.00	B1 V	915 GAM	PER	3 1+9 53 21	2.94	1.16	0.70	0.61	1.06	GRIII+A3:	
255 UPS	2 CAS	0 54+2 58 58	4.64	1.66	0.96	0.77	1.27	G8 III-IV	919 TAU 3	ERI	3 0 6	-23 47	4.07	0.17	0.13	0.22	A4 V	
259 MU	AND	0 54+5 38 17	3.87	0.27	0.12	0.15	0.23	A4 III	921 RHO	PER	3 2+6 38 41	3.39	3.38	1.67	1.78	3.39	M4 IIII	
271 ETA	AND	0 55+1 23 13	4.39	1.64	0.95	0.73	1.21	G8 III-IV	922 1	DEL	ARI	3 9+3 19 35	4.58	2.24	1.49	1.05	1.81	A3 V
280 ALF	SCL	0 56+7 -29 34	4.27	-0.70	-0.16	-0.03	-0.16	B8 IIIP	923 916	BET	PER	3 5+6 40 48	2.15	-0.44	-0.06	0.07	0.77	BB V
285 DEL	CET	1 2+9 86 3	4.27	2.54	1.21	0.88	1.48	K2 III	927 1	OMI	PER	3 6+2 49 28	6.07	0.77	0.59	0.54	0.83	GO V
294 EPS	PSC	1 0+9 41 24	4.78	1.67	0.97	0.78	1.31	K2 III	941 KAP	PER	3 6+8 44 42	3.81	1.81	0.98	0.75	1.25	KO III	
322 D BET	PHE	1 4+3 -46 56	(3.30)	-	0.89	0.71	1.23	G8 III	947 OHG	PER	3 8+7 39 28	4.64	2.13	1.11	0.83	1.39	M4 IIII	
334 ETA	CET	1 6+0 -5 89	5.24	3.43	2.35	1.15	0.84	K3 III	951 951	DEL	ARI	3 9+3 19 35	4.37	1.89	1.00	0.77	1.28	K2 III
335 D PHI	AND	1 7+2 47 1	4.25	-0.41	-0.07	0.09	-0.09	B7 V	953 963	ALF	FOR	3 21+5 49 43	3.85	0.57	0.52	0.66	0.77	F8 IV
337 BET	CET	1 7+5 35 29	2.04	3.53	1.57	1.24	2.24	M0 III	972 ZET	ARI	3 12+6 20 54	4.89	-0.03	-0.02	0.08	0.05	A0 IV	
343 THE	CAS	1 8+6 54 56	4.16	-0.30	0.17	0.19	0.26	A7 V	984 ZET	ERI	3 13+9 48 50	4.80	0.32	0.23	0.23	0.34	A7m	
351 CHI	PSC	1 9+3 20 49	4.66	1.84	1.03	0.76	1.30	G8 III	985 985		3 16+4 65 31	4.84	-0.92	-0.14	-0.02	-0.14	B2 Ve	
352 TAU	1	9+5 23 53	4.51	2.10	1.10	0.82	1.40	K0 III-IV	991 991		3 16+2 34 54	4.82	3.06	1.49	1.05	1.81	K2 II	
360 PHI	PSC	1 11+5 24 46	4.66	1.87	1.03	0.75	1.28	K0 III	992 992	BET	PER	3 5+6 40 48	2.15	-0.44	-0.06	0.07	0.77	K2 III
370 NU	PHE	1 13+3 -45 45	(4.95)	-	0.59	0.49	0.85	F8 V	999 999	OMI	TAU	3 22+7 8 53	2.94	1.16	0.70	0.67	1.12	G5 V
382 PHI	CAS	1 17+5 19 49	4.49	1.18	0.68	0.63	1.20	F4 V	1002 1002	DEL	ARI	3 24+8 43 11	4.94	0.11	0.06	0.07	0.08	A3 V
383 UPS	PSC	1 17+3 27 3	4.76	0.13	0.03	0.08	0.13	A3 V	1003 TAU	ERI	3 17+7 -21 54	4.70	3.43	1.62	1.57	3.04	M3	
390 XI	AND	1 20+0 45 19	4.70	2.05	1.08	0.81	1.34	K0 III-IV	1008 E	ERI	3 18+2 43 43	(4.26)	*	0.72	0.52	1.02	F5 b	
399 PSI	CAS	1 23+1 57 59	4.74	1.97	1.03	0.77	1.29	K0 III	1017 1017	ALF	PER	3 21+5 49 43	4.10	2.88	0.48	0.45	0.78	H5 b
402 THE	CET	1 24+0 57 23	3.59	1.99	1.06	0.76	1.33	K0 III	1030 1030	OMI	TAU	3 22+7 8 53	3.60	1.51	0.90	0.67	1.12	G5 III
403 V DEL	CAS	1 20+2 60 2	4.79	-0.28	-0.13	0.15	0.24	A5 V	1034 1034	DEL	ARI	3 25+2 48 55	4.50	-0.69	-0.13	0.00	-0.09	B3 V
417 CMG	AND	1 25+3 45 12	4.83	0.41	0.42	0.41	0.54	F4 IV	1035 1035	SIG	PER	3 27+7 47 52	4.37	2.89	1.35	1.07	1.81	K3 III
420v ALF	UMI	1 55+7 89 5	2.04	0.97	0.59	0.50	0.81	F8 V	1038 XI	TAU	3 25+0 59 48	4.20	0.18	0.41	0.37	0.76	B9 1a	
429 V GAM	PHE	1 26+6 -63 31	(3.40)	-	1.56	1.26	2.24	K5 III	1040 1040	TAU	3 26+7 9 56	4.74	0.45	0.46	0.56	0.55	I-100	
434 MU	PSC	1 28+0 57 43	2.95	1.38	1.06	1.03	1.60	K4 III	1044 D	PER	3 26+5 49 22	4.68	-0.68	-0.10	0.01	-0.09	B9 V	
437 ETA	PSC	1 29+3 15 8	3.62	1.59	0.97	0.71	1.20	G8 III	1046 D		3 26+9 55 19	5.13	0.08	0.01	0.09	0.01	A1 V	
440 DEL	PHE	1 29+6 -49 17	(3.94)	-	0.98	0.75	1.26	K0 III-IV	1052 SIG	PER	3 27+7 47 52	4.37	2.89	1.35	1.07	1.81	K3 III	
442 CHI	CAS	1 31+3 51 46	4.22	1.72	1.16	1.00	1.77	K1 V	1056 5	TAU	3 28+7 12 48	4.09	2.16					

## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

B.S.	NAME	R.A.(1960) DEC.	V	U-V	B-V	V-R	V-I	MK SP.	B.S.	NAME	R.A.(1960) DEC.	V	U-V	B-V	V-R	V-I	MK SP.
1240	TAU 9 ERI	3 58+3 -24 7	4.66	-0.53	-0.14	0.02	-0.10	A p	1791	BET TAU	5 23+8 28 35	1.66	-0.62	-0.13	-0.01	-0.09	B7 II
1242	OMG 1 ERI	4 1.1 59 4	5.07	0.99	0.49	0.51	0.90	Fo II	1810	I1A TAU	5 25+2 21 54	4.89	-0.91	-0.14	-0.04	-0.19	B3 V
1251	NU TAU	4 1.0 5 53	3.91	0.08	0.03	0.09	0.09	A1 V	1811 D	PSI ORI	5 24+7 3	4.61	-1.14	-0.22	-0.08	-0.31	B2 IV
1256	37 TAU	4 2+3 21 58	4.37	2.02	1.07	0.80	1.33	KO III	1829	BET LEP	5 26+5 -20 47	2.81	-1.31	0.82	0.65	1.09	G5 III
1261	LAM PER	4 3+6 50 15	4.29	-0.02	0.02	0.09	0.11	B9 V	1839 D	32 ORI	5 28+6 5 55	4.20	-0.69	-0.13	-0.01	-0.17	B4 IV
1273	48 PER	4 5+8 47 36	4.04	-0.50	-0.06	0.13	0.11	B3 Vp	1843	CH1 AUR	5 30+1 32 10	4.77	-0.09	0.35	0.37	0.44	B5 Iab
1298	OMG 1 ERI	4 9.9 -6 56 4.05	0.47	0.33	0.32	0.48	0.22	I1-II-III	1845 V	I19 TAU	5 29+6 18 34	4.35	-2.04	1.75	3.20	M2 Ib	
1302	DEL HOR	4 7.5 -42 6 14.921	1.02	0.25	0.35	0.56	0.0	Fo	1853 S	HD 36395	5 29+5 -3 41	7.97	-2.68	1.47	1.44	2.58	M1 V
1303	HU PER	4 12.0 48 19	4.14	1.59	0.95	0.77	1.31	GO Ib	1852 V	DEL ORI	5 30+0 -20 2	2.21	-1.27	-0.21	-0.07	-0.28	095 II
1306	52 PER	4 12.2 40 23	4.71	1.66	1.01	0.79	1.35	G5 Ib+A2	1855 S	UPS ORI	5 30+0 -7 21	4.63	-1.33	-0.26	-0.06	-0.32	80 V
1309	46 TAU	4 11.4 7 37	5.28	0.37	0.36	0.36	0.55	F3 V	1861	CH1 AUR	5 30+1 -6 13	5.35	-1.12	-0.19	-0.04	-0.22	B1 V
1311 D	TAU 4 11.8 9 10	4.83	1.32	0.80	0.64	1.10	I05	1862	EPS COL	5 29+8 -35 31	(3.86)	-1.12	0.82	1.42	K0		
1318 D	39 ERI	4 12.5 -10 21 4.85	2.33	1.17	0.85	1.42	K2 III+G2V	1865	ALF LEP	5 31+0 -17 51	2.58	0.47	0.19	0.22	F0 Ib		
1320	HU TAU	4 13.4 8 48	4.30	-0.56	-0.05	0.06	0.00	B3 V	1868 V	ORI	5 31+4 -1 12	5.35	-1.09	-0.19	-0.07	-0.26	B1 V
1324	B PER	4 15.2 50 12	4.65	0.10	0.05	0.15	0.19	A2	1876	PHI ORI	5 32+6 9 28	4.44	-1.12	0.05	0.01	-0.16	B1 V
1325	OMH 2 ERI	4 13.5 -7 40	4.42	1.27	0.82	0.68	1.13	K1 V	1877/80	LAH ORI	5 32+7 9 54	3.39	-1.17	-0.18	-0.04	-0.20	88
1326	ALF HOR	4 12.7 -42 24 13.851	1.11	0.86	1.45	K1 III	1887	CH1 V	5 33+0 -2 24	4.77	-1.25	-0.23	-0.11	-0.36	80 V		
1327	DEL TAU	4 16.9 65 3	5.28	1.31	0.81	0.65	1.08	M5 III	1892 D	42 ORI	5 33+4 -5 52	4.60	-1.13	-0.19	-0.06	-0.26	B2 III
1329	OMG TAU	4 14.9 20 29	4.93	0.34	0.25	0.25	0.36	A m	1893	THE ORI	5 33+3 -1 25	4.72	-0.86	0.00	0.33	0.56	B0.5 V
1346	GAM TAU	4 17.5 15 32	3.64	1.77	0.99	0.73	1.20	KO III	1895	THE 1 ORI	5 33+3 -5 25	5.13	-0.95	-0.01	0.23	0.41	05P
1347	41 ERI	4 16+4 -33 54 (3.55)	*	-0.11	-0.01	-0.12	A p	1896	THE ORI	5 33+3 -5 25	6.70	-0.74	0.08	0.31	0.51	09.5 VI	
1350	53 PER	4 18.6 46 25	4.86	-0.56	-0.03	0.09	0.04	B6 III	1897	THE 2 ORI	5 33+3 -5 25	5.07	-1.03	0.08	0.15	0.16	09.5 VP
1373	DEL TAU	4 20.6 17 27	3.76	1.82	0.99	0.73	1.20	K1 III	1899 D	IOT ORI	5 33+5 -5 56	2.75	-1.30	-0.23	-0.06	-0.28	09 III
1380	-64 TAU	4 21.7 17 21	4.81	0.27	0.15	0.17	0.24	A7.5 V	1901 W	ORI	5 33+3 -4 53	5.25	0.43	0.25	0.28	0.42	8f0
1387	KAP TAU	4 23+0 22 12	4.23	0.27	0.12	0.16	0.21	A7 V	1903 V	EPS ORI	5 34+2 -1 14	1.70	-1.21	-0.19	-0.01	-0.19	B8 Ia
1389	68 TAU	4 23+2 17 50	4.29	0.12	0.04	0.09	0.10	A3 V	1907	PHI 2 ORI	5 34+7 9 16	4.08	1.64	0.95	0.75	1.31	G8 IIIP
1392	UPS TAU	4 23+9 22 43	4.29	0.39	0.25	0.27	0.41	A8 Vn	1908	CH1 V	5 34+8 11 1	5.90	3.53	1.60	2.13	K5 II	
1393	43 ERI	4 22.5 -34 7 (3.94)	1.48	1.17	0.80	2.00	M1 III	1910	ZET TAU	5 35+3 21 7	3.03	-0.18	-0.08	-0.09	-0.26	B2 IVp	
1394	71 TAU	4 24.0 15 31	4.31	0.41	0.25	0.27	0.42	Fo V	1931 D	SIG ORI	5 36+7 -2 37	3.83	-1.26	-0.24	-0.04	-0.26	09.5 V
1396	P1 TAU	4 24+3 14 37	4.70	1.71	0.98	0.72	1.23	G8 III	1934	OMG ORI	5 37+1 4 6	4.59	-0.88	-0.11	0.05	-0.05	B3 III
1409	EPS TAU	4 26+3 19 6	3.53	1.89	1.01	0.73	1.23	K1 III	1937	49 ORI	5 36+9 -7 14	4.81	0.23	0.14	0.15	0.21	A4 IV
1411	THE 1 TAU	4 26+2 15 52	3.85	1.63	0.93	0.71	1.17	G9 III	1946 D	126 TAU	5 39+0 16 31	4.85	-0.76	-0.12	-0.01	-0.13	B3 IV
1412	THE 2 TAU	4 26+4 15 47	3.41	0.34	0.18	0.19	0.28	A7 IVN	1948/9	ZET ORI	5 38+7 -1 58	4.75	-1.26	-0.21	-0.06	-0.27	09.5 Ib
1414 *	79 TAU	4 26+6 12 58	5.00	0.34	0.22	0.20	0.28	d46a	1956	ALF COL	5 38+2 -3 46	(2.63)	-0.15	-0.01	-0.11	-0.11	B2 Ve
1427	P1 TAU	4 28+2 16 7	4.78	0.32	0.17	0.17	0.24	A6 Vn	1953	PI	5 40+4 1 28	4.89	-2.24	1.18	0.88	1.49	K0 III
1437	45 ERI	4 29+9 -8 0	4.20	2.76	1.32	0.94	1.61	K3 II-III	1983	GAM LEP	5 42+8 -22 20	3.58	0.51	0.48	0.45	0.72	F6 V
1444	RHO TAU	4 31.6 14 46	4.63	0.33	0.25	0.24	0.36	A6 Vn	1995	TAU AUR	5 44+4 39 11	4.52	1.84	0.94	0.72	1.22	G3 III
1453	UPS 1 ERI	4 32+0 -29 51	4.50	1.72	0.99	0.75	1.29	G6	1998	ZET LEP	5 45+1 -14 50	3.56	0.18	0.08	0.13	0.16	A3 V
1454	58 PER	4 35+9 41 11	4.27	2.02	1.22	0.97	1.65	G5 II-II+A	2004	KAP ORI	5 45+9 -9 41	2.06	-1.18	-0.18	-0.03	-0.21	8.5 P. Ia
1457	ALF TAU	4 33+6 16 26	3.64	3.50	1.95	1.22	2.15	Tau 2010	2010	134 TAU	5 47+3 12 38	4.90	-0.21	-0.07	0.03	-0.03	B9 IV
1458	88 TAU	4 33+5 10 4	4.26	0.30	0.18	0.19	0.28	A p	2011	UPS AUR	5 48+3 37 17	4.73	3.55	1.62	1.34	2.42	E1
1463	NU ERI	4 34+3 -3 26	3.92	-1.09	-0.21	-0.19	-0.29	B2 III	2012	NU	5 48+7 39 8	3.97	2.24	1.13	0.82	1.38	K0 III
1464	UPS 2 ERI	4 34+0 -30 39 (3.81)	*	0.97	0.75	1.24	KD III	2018	CH1 V	5 48+8 32 7	6.22	3.83	1.76	2.94	M3 III		
1473	90 TAU	4 35+9 12 26	4.27	0.26	0.13	0.13	0.19	A6 Vn	2027	XI V	5 51+5 55 45	4.98	0.17	0.05	0.09	0.12	A2p
1479	SIG 2 TAU	4 37+0 15 30	4.69	0.31	0.16	0.17	0.25	A5 V	2034	136 TAU	5 50+8 27 36	4.61	-0.01	-0.02	0.04	0.05	B9.5 V
1481 D	53 ERI	4 36+3 -14 23	3.85	2.11	1.09	0.82	1.39	K2 III	2035 DEL	LEP	5 49+6 -20 53	3.79	1.70	0.98	0.79	1.36	G8+ III
1496 V	54 ERI	4 38+7 -19 45	4.33	3.40	1.61	1.55	2.99	F4	2040	BET COL	5 49+9 -35 47	(3.10)	1.15	0.85	1.43	K2 III	
1497	TAU TAU	4 39+8 22 53	4.29	-0.71	-0.14	-0.03	-0.15	B2 V	2047	CH1 1 ORI	5 52+1 20 17	4.40	0.69	0.60	0.52	0.83	GO V
1502	ALF CAE	4 39+3 -41 56	4.441	*	0.36	0.34	0.55	B2 V	2056	LAH COL	5 51+7 -39 49	(4.88)	-0.17	-0.06	-0.22	85 V	
1520	MU ERI	4 39+5 -3 20	4.02	-0.76	-0.16	-0.05	-0.19	B5 IV	2061 V	ALF ORI	5 53+0 7 24	6.69	-1.92	1.86	1.60	2.07	M2 Iab
1542	ALF CAM	4 50+1 66 17	4.27	-0.85	0.07	0.12	0.12	O9.5 Ia	2077 DEL	AUR	5 56+3 54 17	1.73	1.90	1.00	0.75	1.25	K0 III
1543	P1 ORI	4 47.7 6 54	3.19	0.45	0.46	0.42	0.68	F6 V	2084	139 TAU	5 55+5 25 58	4.83	-0.99	-0.07	0.05	-0.05	B1 B
1544	P1 2 ORI	4 48.5 8 51	4.51	4.35	0.04	0.01	0.06	A0 V	2085	ETA LEP	5 56+6 -14 30	3.69	0.34	0.32	0.31	0.48	F0 IV
1547 *	97 TAU	4 49.0 18 47	5.10	5.10	0.14	0.22	0.34	J45	2088 V	BET AUR	5 56+5 47 57	4.60	-0.09	0.02	0.09	0.08	A2 V
1552	P1 4 ORI	4 49.1 9 32	3.67	-0.96	-0.15	-0.02	-0.17	B2 IV	2091 V	P1 AUR	5 57+0 45 57	4.25	3.55	1.72	1.68	3.16	M3.5 II
1560	OMG ERI	4 50+9 -5 31	4.39	0.40	0.23	0.32	0.47	A9 IV	2095 D	THE AUR	5 57+0 37 13	2.63	-0.28	-0.08	0.01	-0.04	B9.50 V
1567 V	P1 5 ORI	4 52.2 23	3.73	-1.01	-0.19	-0.05	-0.24	B2 III	2106 GAM	COL	5 56+1 -35 17 (4.35)	-0.17	-0.11	-0.09	-0.23	3.19	B5 V
1568 D	7 CAM	4 54+1 53 42	4.47	-0.05	-0.02	0.10	0.09	A1 V	2113	5 58+0 -3 29	6.52	2.43	1.22	2.23	1.60	K2 III	
1570	P1 1 ORI	4 54.7 10 6	4.66	0.16	0.08	0.11	0.14	A0 V	2120	ETA COL	5 57+9 -42 49 (3.94)	1.94	1.32	1.02	1.40	1.40	K0 III
1577	IOT AUR	4 54+4 33 6	2.67	3.29	1.56	1.06	1.28	K2 III	2120 D	MU ORI	6 0+2 9 40	4.14	-0.30	0.16	0.20	0.27	A
1580	OMH 2 TAU	4 54+1 13 27	4.06	2.26	1.15	0.88	1.50	K2 III	2128	3 MON	6 12+9 -6 16	3.96	2.74				

## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

SP.	B.S.	NAME	R.A. (1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.	B.S.	NAME	R.A. (1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.				
III	V	XI	GEM	6 43.0	-12 56	3.34	0.47	0.42	0.40	0.63	F5 IV	3282	PUP	8 19.8	-32 56	(4.82)	+	1.44	1.01	1.75	H0 III			
IV	V	1491	ALF	CMA	6 43.4	-16 40	-1.45	-0.03	-0.01	0.00	-0.03	A1 V	3294	D	B VEL	8 21.2	-48 21	(4.82)	-	-0.16	-0.04	-0.16	B1 V	
III	V	1506	IB	MON	6 45.8	-2 27	4.43	2.16	1.11	0.78	1.33	K0 III	3314	P	B VEL	8 23.7	-3 46	1.89	-0.08	-0.04	0.03	-0.02	A0 V	
IV	V	2527			6 54.3	77 2	4.54	3.04	1.36	1.02	1.72	K4 III	3323	OHI	U MA	8 27.0	60 51	3.36	1.38	0.85	0.68	1.10	G4 II-III	
IV	V	2538 *	KAP	CMA	6 48.3	-32 28	(3.96)	-	-0.24	-0.07	-0.27	B2 Ve	3403	PI 2	U MA	8 36.7	64 28	A.61	2.34	1.17	0.88	1.50	K2 III	
Tab															C	VEL	8 33.5	-49 49	(5.00)	+	1.31	0.97	1.65	K0
IV	V	2540 D	THE	GEM	6 50.2	34 1	3.61	0.25	0.11	0.11	0.18	A3 III	3407		9	HYA	8 39.2	-15 48	4.87	2.02	1.08	0.81	1.34	K1 III
V	V	1560 D	15	LYN	6 53.6	58 29	4.35	1.39	0.85	0.66	1.09	G5 III-IV	3410	DEL	HYA	8 35.5	-5 51	4.17	0.02	0.00	0.04	0.04	A0 V	
5 II	V	1564 D	3B	GEM	6 52.4	13 14	4.65	0.38	0.31	0.32	0.52	F0 Vp	3418	SIG	HYA	8 36.7	3 29	4.43	2.48	1.20	0.87	1.43	K2 III	
V	V	1571	15	CMA	6 51.8	-20 10	4.82	-1.17	-0.21	-0.11	-0.33	B1 IV	3426	t	VEL	8 36.2	-42 51	(4.13)	+	0.11	0.16	0.32	A9 II	
V	V	1574	THE	CMA	6 52.3	-11 59	4.07	3.12	1.44	1.10	1.89	K3+ III	3438	BET	PYX	8 38.5	-35 10	(1.98)	+	0.92	0.67	1.15	G3 III	
V	V	2580	OMI 1	CMA	6 52.5	-25 8	3.92	3.74	1.72	1.51	1.97	K3 IAB	3441	9	HYA	8 39.2	-15 48	4.87	2.02	1.08	0.81	1.34	K1 III	
IVb	V	2583	16	LYN	6 54.7	45 5	4.90	0.07	0.05	0.05	0.07	A4 V	3445	b	VEL	8 39.3	-46 30	(3.82)	+	0.71	0.64	1.26	F2 Ia	
IVb	V	2590 D	PI	CMA	6 53.9	-20 5	4.67	0.42	0.36	0.37	0.57	F2	3449	GAM	CNC	8 41.0	21 37	4.66	0.03	0.02	0.07	0.07	A1 Y	
IV	V	2596	IOT	CMA	6 54.4	-18 59	4.39	-0.74	-0.06	0.05	-0.06	B3 II	3452	t	VEL	8 39.9	-47 11	(4.76)	+	0.11	0.21	0.41	A3 II	
V	V	2608			6 55.2	-46 40	3.94	1.69	1.46	2.66	H3	3454 S	ETA	HYA	8 41.1	3 33	4.30	-0.94	-0.20	0.06	-0.25	B3 V		
V	V	2618	EPS	CMA	6 57.1	-28 35	1.50	-1.13	-0.21	-0.09	-0.29	B2 II	3459	31	HOM	8 41.7	7 5	4.60	1.34	0.84	0.65	1.09	G2 Ib	
V	V	2648	SIG	CHM	7 0.1	23 53	3.43	3.59	1.72	1.32	2.32	MO IBD	3461	DEL	CNC	8 42.4	18 18	3.93	2.08	1.08	0.78	1.32	K0 III	
5 V	V	2649	19	HOM	7 0.9	-4 11	5.01	-1.12	-0.20	-0.06	-0.24	B1 V	3468	ALF	PYX	8 42.0	-33 2	(3.70)	-	-0.21	-0.07	-0.23	B2 II	
V	V	2650 V	ZET	GEM	7 1.7	10 20	4.65	1.15	1.74	0.97	1.17	G1.5 Ib	3475	IOT	CNC	8 43.3	28 54	4.02	1.81	1.03	0.74	1.26	G8 II	
V	V	2653	OH1 2	CMA	7 1.4	-23 46	3.01	-0.94	-0.11	0.01	-0.08	B3 Ia	3477	t	VEL	8 42.9	-42 30	(4.06)	+	0.87	0.64	1.13	igG5	
5 V:	V	2657	GAM	CMA	7 1.9	-15 34	4.11	-0.57	-0.11	-0.01	-0.11	B8 II	3482 D	EPS	HYA	8 44.4	6 34	3.39	1.05	0.68	0.61	0.99	G0 COMP.	
5 VP	V	2653	DEL	CMA	7 6.8	-26 20	1.80	1.20	0.67	0.52	0.85	F0 V	3484	12	HYA	8 44.5	-13 24	4.32	1.30	0.90	0.70	1.16	G8 III	
V	V	2667	TAU	GEM	7 8.6	30 20	4.42	2.67	1.26	0.94	1.57	K2 III	3485 D	DEL	VEL	8 43.6	-24 36	(1.96)	+	0.04	0.05	0.09	A0: V	
V	V	2701 *	20	HOM	7 8.2	-4 10	4.91	1.82	1.03	0.78	1.31	K0 III	3487		HYA	8 44.6	-45 53	(3.90)	+	0.01	0.06	0.19	A0 III	
V	V	2702 A	PU	T	7 7.5	-39 33	(4.82)	-	-0.20	-0.07	-0.22	B3 V	3492	RHO	HYA	8 46.3	6 0	4.37	-0.10	-0.05	0.05	-0.02	A0 V	
III	V	2714	DEL	MON	7 9.8	-0 26	4.15	0.03	0.00	0.05	0.09	A1 IV	3518	GAM	PYX	8 48.8	-27 34	4.00	2.65	1.27	0.95	1.61	K4 III	
IV	V	2720	1	PU	7 11.4	-46 42	(4.48)	-	0.36	0.32	0.54	F0 V	3527	t	VEL	8 49.2	-46 22	(5.10)	+	-0.21	-0.06	-0.23	B0 III	
5 V	V	2745 V	27	CMA	7 12.6	-26 17	4.65	-0.85	-0.16	-0.02	-0.03	B3 II	3547	ZET	HYA	8 53.3	6 6	3.10	1.81	1.01	0.72	1.21	G8 III	
V	V	2748 V	L	PU	7 12.3	-44 35	5.09	-	1.55	2.72	5.09	M5e	3556	DEL	PYX	8 53.8	-27 32	4.87	0.27	0.12	0.14	0.19	A3 V	
V	V	2749	DMG	CMA	7 13.2	-26 42	3.82	-0.88	-0.14	0.03	-0.07	B3 IVE	3569	IOT	UMA	8 56.5	48 12	3.15	0.27	0.19	0.21	0.29	A7 V	
IV	V	2751			7 15.5	49 32	5.04	0.17	0.08	0.15	0.21	A3 III-IV	3572	ALF	CNC	8 56.3	12 1	4.26	0.28	0.13	0.16	0.20	A m	
IV	V	2762			7 13.6	-48 13	(4.75)	-	0.10	-0.01	-0.09	B8 II	3576	RHO	UMA	8 59.0	47 47	4.76	3.42	1.54	1.44	2.70	M3 IIIb	
5 Tb	V	2763	LAM	GEM	7 15.8	16 37	3.59	0.21	0.11	0.12	0.17	A3 V	3579		HYA	8 58.0	-41 57	3.97	0.48	0.43	0.41	0.62	F5 V	
V	V	2764			7 14.9	-23 15	4.78	3.58	1.69	1.28	2.44	2M0	3591	t	VEL	8 58.6	-41 16	(4.45)	+	0.66	0.56	0.96	F8 III	
V	V	2765			7 15.0	-27 49	4.60	3.53	1.60	1.17	2.73	H3	3594	KAP	UMA	9 0.9	-0.49	1.79	1.61	0.01	0.09	0.09	B9 B	
V	V	2773	PI	PU	7 15.7	-37 1	(2.70)	-	1.52	1.24	2.15	K5 III	3612		9	4.0	38 37	4.56	1.85	1.04	0.74	1.24	GB Ib-II	
V	V	2777 D	DEL	GEM	7 17.7	22 3	3.52	0.40	0.34	0.35	0.53	F0 IV	3614		1	2.8	-46 56	(3.74)	-	1.20	0.82	1.42	K2 III	
5 Ta	V	2781 V	29	CMA	7 17.0	-24 30	4.75	-1.15	-0.15	-0.09	-0.21	071	3616 D	SIG 2	UMA	9 6.9	67 18	4.83	0.54	0.50	0.45	0.73	FT- IV-V	
V	V	2782	TAU	CMA	7 17.1	-24 53	4.48	-1.11	-0.19	-0.03	-0.15	B9 III	3619	15	UMA	9 6.1	51 46	4.48	0.38	0.27	0.28	0.37	A m	
V	V	2787			7 16.9	-36 40	(4.61)	-	-0.08	0.10	0.05	B3 ve	3624	TAU	UMA	9 7.7	63 41	4.69	0.51	0.35	0.47	0.47	A7 m	
V	V	2812			7 20.4	-18 57	4.96	-0.43	-0.06	0.06	0.03	B9	3628	KAP	PYX	9 6.3	-25 42	4.56	3.47	1.59	1.23	2.14	EMO	
V	V	2818	ZI	LYN	7 23.7	49 18	4.63	-0.04	-0.02	-0.02	-0.01	A1 V	3634	LAM	VEL	9 6.5	-42 16	(2.40)	1.62	1.24	1.24	1.19	K5 Ib	
5 V	V	2821	IOT	ETA	7 23.2	27 53	3.79	1.49	1.03	0.76	1.27	K0 III	3634		UMA	9 6.5	-44 43	(4.99)	+	0.25	0.30	0.52	B3 Ia	
V	V	2827	ETA	CMA	7 22.5	-29 13	4.48	2.41	0.77	0.07	0.02	B2 V	3682	18	UMA	9 13.3	54 11	4.04	0.25	0.17	0.26	0.36	A5 V	
V	V	2845	BET	CMA	7 25.0	-36 40	2.90	-0.37	-0.09	0.03	-0.04	B7 V	3685	THE	HYA	9 12.3	2 29	3.88	-0.19	-0.07	0.01	-0.05	89.5 vp	
III	V	2852 S	PHO	GEM	7 26.5	31 52	4.17	-0.51	0.40	0.38	0.61	F3 III	3749	1	VEL	9 18.1	-38 24	(4.93)	+	1.11	0.80	1.37	K2 Z	
V	V	2854	GAM	CMA	7 26.0	9 1	4.29	2.97	1.43	1.07	1.87	K3 III	3684		9	14.1	-37 15	(4.62)	+	0.47	0.39	0.67	F3 IV-V	
V	V	2856	6	CMA	7 27.6	-12 16	4.54	2.67	1.29	0.88	1.52	K2 III	3690 D	38	LYN	9 15.3	36 59	3.83	0.10	0.05	0.14	0.17	A2 V	
V	V	2874			7 28.1	-22 56	4.83	0.45	0.25	0.26	0.53	A3 Ib	3705	ALF	CNC	9 18.0	34 34	3.14	1.55	1.21	1.21	GB III		
V	V	2881	ZET	GEM	7 29.1	-34 52	(4.64)	-	0.51	0.40	0.78	F3 III	3709	27	HYA	9 18.5	-9 23	4.79	1.62	0.94	0.68	1.12	GB III	
V	V	2897	1	PU	7 30.9	-18 48	4.51	-1.08	-0.06	0.15	0.13	BD Vpe	3718	THE	PYX	9 19.7	-25 48	4.72	3.69	1.54	1.39	2.47	M1 III	
V	V	2903	OMI	CMA	7 31.4	45 45	3.45	3.49	1.54	1.22	2.14	K5 V	3731	KAP	HYA	9 22.3	26 21	4.46	2.34	1.23	0.90	1.52	K2 III	
V	V	2904	1	PU	7 31.7	-20 17	4.56	0.15	0.08	0.15	0.22	A3 V												

## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

B.S.	NAME	R.A. (1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.	B.S.	NAME	R.A. (1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.	
4092	HU	10 23+7	-6 51	5.58	3.40	1.51	1.25	2.19	MO III	4942	XI 2	CEN 13	4+6	-49 42	(4+26)	-	-0.20	-0.08	-0.28	B2 V
4094	HU	HYA 10 24+4	-16 50	3.29	1.48	1.11	1.04	K4 III	4954	A1	COM 13	5+3	27 51	4+82	3-32	1+45	1+18	1+99	K5 III	
4100 D	BET	LU 10 25+6	-36 53	4.20	1.59	0.90	0.69	1.15	G8 III-IV	4963 D	THE	VIR 13	7+9	-50 20	4+37	-0.00	-0.01	0.04	0.04	A1 V
4104	ALF	ANT 10 25+3	-30 52	4.24	*	1.43	1.10	1.89	MO III	4979	13	9+8	-37 35	(4+84)	*	0.69	0.37	0.94	dG3	
4112	36	UMA 10 28+1	56 11	4.84	0.50	0.52	0.48	0.75	FB V	4983	BET	COM 13	10+0	28 5	4+26	0.67	0.58	0+8	0.77	GO V
4119	BET	SEX 10 28+3	-0 27	5.09	-0.67	-0.13	-0.03	-0.18	B6 V	5017	20	CVN 13	15+8	40 47	4+73	0+50	0+30	0+25	0+40	F0 III-IV
4132	HU	10 30+9	40 37	4.74	0.31	0.23	0.19	0.28	A7 IV	5019	61	VIR 13	16+3	-18 5	4+74	0+98	0+71	0.58	0+94	G6 V
4133	RHO	LEO 10 30+7	9 31	3.85	-1.09	-0.13	-0.04	-0.21	BL I	5020	GAM	HYA 13	16+7	-22 58	2+98	1+60	0+92	0.63	1+09	G6 III
4163 V	U	HYA 10 35+6	-13 11	4.92	8.48	2.75	1.80	3.08	C73	5026	13	18+2	-52 32	5+48	*	-0.14	-0.08	-0.23	B5 III	
4166	37	LMI 10 36+5	32 11	4.70	1.37	0.81	0.65	1.05	K3 III	5028	IOT	CEN 13	18+3	-36 10	2+73	0.05	0.02	0.06	0.05	A2 V
4167 D	P	VEL 10 35+6	-48 1	(3.04)	*	0.31	0.26	0.42	FOP	5054	ZET A	LMA 13	22+3	55 8	2+27	0+01	0+02	0+06	0+03	A2 V
4216 D	HU	VEL 10 45+1	-49 13	(2.68)	*	0.88	0.69	1.18	G5 III	5057	ZET B	LMA 13	22+3	55 8	3+95	0+22	0+13	0+13	0+18	A M
4232	NU	HYA 10 47+6	-15 59	3.10	2.51	1.74	0.91	1.55	K3 III	5056 V	ALF	VIR 13	23+1	-10 57	0+95	-1+18	-0+25	-0+09	-0+32	B1 V
4247	46	LMI 10 51+1	34 26	3.82	1.98	1.04	0.83	1.37	K0 III-IV	5062	BO	UMA 13	23+6	55 11	4+03	0+15	0+18	0+24	0+24	A5 V
4248	OMG	UMA 10 51+7	43 24	4.71	-0.11	-0.05	0.07	0.03	A1 V	5068	69	VIR 13	25+3	-15 46	4+75	2+16	1+10	0.79	1+32	K1 III
4259/50	54	LEO 10 53+5	28 58	4.32	0.03	0.02	0.08	0.09	A1 V	5072	70	VIR 13	26+5	14 0	4+97	0+99	0+71	0+61	0+97	G5 V
4273	IOT	ANT 10 54+8	-35 55	(4.58)	*	1.02	0.75	1.28	G5 III	5080 V	R	HYA 13	27+5	-23 5	5+98	3+31	1+52	2+99	5+37	H7e
4287	ALF	CRT 10 57+8	-10 5	4.56	2.06	1.07	0.78	1.33	K0 III	5089 D	CEN	13	28+7	-39 13	3+88	2+21	1+18	0.84	1+43	G8 III
4293	58	UMA 10 58+3	-42 1	(4.39)	*	0.12	0.14	0.21	A2 IV	5095	74	VIR 13	29+9	3 46	3+56	1+60	1+44	2+61	M2+ III	
4295	BET	UMA 10 59+4	53 56	2.38	-0.04	-0.01	0.08	0.04	A1 V	5105	78	VIR 13	32+1	3 52	4+94	0+02	0.02	0+06	0+04	A p
4299	61	LEO 10 59+8	-2 16	4.74	3.54	1.62	1.30	2.27	K5 III	5107	ZET	VIR 13	32+7	-0 24	3+38	0+20	0+12	0+07	0+13	A3 Vm
4300	60	LEO 11 0+2	20 24	4.61	0.08	0.04	0.08	0.06	A1 V	5110	13	33+0	-37 23	5+01	0+45	0+39	0+41	0+71	F2 IV	
4301	ALF	UMA 11 1+3	61 58	1.80	2.01	1.07	0.81	1.39	K0 III	5112	24	CVN 13	32+3	55 5	3+95	0+23	0+12	0+13	A4 V	
4310	CHI	LEO 11 3+0	7 33	4.62	0.41	0.33	0.32	0.49	F2 III-IV	5127 D	25	CVN 13	35+7	-36 29	4+83	0+32	0+23	0+28	0+40	A7 III
4335	PSI	UMA 11 7+4	44 43	3.01	2.26	1.15	0.82	1.41	K1 III	5132 EPS	CEN	13	37+3	-53 16	(2+29)	*	-0.24	-0+15	-0+40	B1 V
4343	BET	CRT 11 9+7	-22 36	4.48	0.10	0.03	0.09	0.12	A2 III-IV	5154	83	UMA 13	39+2	-54 53	4+65	3+60	1+64	1+39	2+51	M2 III
4357	DEL	LEO 11 12+0	20 45	2.57	0.21	0.11	0.16	0.19	A4 V	5168	1	CEN 13	42+4	-32 31	(4+23)	0+38	0+34	0+55	0+50	F0 V
4359	THE	LEO 11 12+1	15 39	3.34	0.05	-0.01	0.08	0.05	A2 V	5185	TAU	BOO 13	43+4	-17 39	4+49	0.53	0+49	0+40	0+65	F7 V
4362	72	LEO 11 13+1	23 18	4.60	3.51	1.66	1.49	2+80	E2	5190	N	CEN 13	47+1	-41 29	(3+40)	*	-0.22	-0+11	-0+33	B2 IV
4368	PHI	LEO 11 14+6	-3 28	4.47	0.36	0.21	0.27	0.38	A7 III-IV	5191	ETA	LMA 13	46+0	49 31	1+86	-0.85	-0+18	-0+07	-0+26	B3 V
4371	T5	LEO 11 15+2	3 16	5.17	3.39	1.51	1.27	2+21	MO III	5192 *	2	CEN 13	47+2	-34 15	4+21	2+92	1+49	2+16	4+01	sgM6
4374/5	XI	UMA 11 15+6	31 46	3.79	0.64	0.59	0.52	0.87	GO V	5193	MU	COM 13	47+2	-42 17	2+97	*	-0.17	0+06	-0+05	B2 Vipne
4377	NU	UMA 11 16+3	33 19	3.49	2.95	1.40	1.06	1.76	K3 III	5200	UPS	BOO 13	47+5	16 0	4+07	3+39	1+52	2+07	K5+ III	
4380	55	UMA 11 17+0	38 24	4.79	0.15	0.12	0.14	0.16	A2 V	5210/1	3	CEN 13	49+3	-32 48	(4+31)	-0.15	-0+04	-0+17	0+05	B5 Vp/S9 V
4382	DEL	CRT 11 17+3	-14 38	3.56	2.10	1.11	0.83	1.43	GB III-IV	5217	13	51+1	-53 11	5+92	*	0.02	0+05	0+06	0+05	B5 V
4386	SIG	LEO 11 19+1	6 15	4.06	-0.19	-0.07	0.01	-0.06	B9 V	5219	13	50+1	-38 38	4+75	3+59	1+64	1+52	2+79	RM2	
4392	56	UMA 11 20+6	43 42	4.98	1.80	0.98	0.72	1.19	GB II	5221 D	4	CEN 13	50+8	-31 44	(4+72)	-0.16	-0+03	-0+16	-0+16	B5 III
4399 D	IOT	LEO 11 21+8	10 45	3.94	0.49	0.43	0+40	0.62	F2 IV	5226	10	DRA 13	50+3	66 55	4+65	3+44	1+58	1+58	2+93	GM3
4405 D	GAM	CRT 11 22+9	-17 28	4.08	0.31	0.21	0.23	0.34	A5 V	5231	ZET	CEN 13	53+0	-47 6	(2+54)	*	-0.24	-0+13	-0+34	B2 IV
4434	LAN	DRA 11 29+1	69 33	3.85	3.59	1.62	1.32	2+31	MO III	5235	ETA	BOO 13	52+8	18 36	2+68	0.81	0+58	0+45	0+74	GO IV
4450	XI	HYA 11 32+6	-31 38	5.04	1.65	0.94	0+70	1+19	G7 III	5248	PHI	CEN 13	55+8	-41 54	(3+82)	*	-0.21	-0+13	-0+34	B2 V
4456 S	9CAB	LEO 11 32+6	17 1	5.95	-0.30	-0.16	-0.07	-0.26	B3 V	5249	UPS 1	CEN 13	56+2	-44 37	(3+86)	*	-0.21	-0+13	-0+35	B2 V
4468	THE	CRT 11 34+7	-9 35	4.70	-0.24	-0.08	0.03	-0.04	B9 V	5250	UPS 2	CEN 13	59+2	-45 24	(4+34)	*	0.57	0+49	0+84	F7 I-II
4471	UPS	LEO 11 34+9	-0 36	4.30	1.75	1.01	0.73	1+25	G9 III	5264	TAU	VIR 13	59+6	1 44	4+27	0.23	0+09	0+14	0+20	A3 III
4494	OMI	HYA 11 3d+2	-34 31	(4.70)	*	-0.07	0.02	-0.03	BB	5265	CHI	CEN 14	3+5	-40 59	(4+35)	*	-0.20	-0+10	-0+33	B2 V
4496	61	UMA 11 38+9	34 26	5.31	1.03	0.78	0.59	0.95	GB V	5267	PI	HYA 14	1+4	-26 29	3+28	2+16	1+12	3+87	1+42	K2 III
4514	ZET	CRT 11 42+7	-18 8	4.73	1.71	0.98	0.71	1+19	G8 III	5268	+THE	CEN 14	4+3	-38 10	(2+05)	*	0+28	0+76	1+47	K2 III-IV
4517	NU	VIR 11 42+8	6 45	4.09	3.28	1.49	1.23	2+25	M1 III	5291	ALF	DRA 14	3+3	65 34	3+66	-0.12	-0.05	-0.03	-0.10	A10 III
4518	CHI	UMA 11 44+0	48 0	3.72	3.22	1.57	1.17	1.87	K0 III	5299	TAU	VIR 14	6+3	64 3	5+27	1+58	1+83	3+51	M4-4.5 III	
4527 D	O	LEO 11 50+2	29 26	4.51	0.55	0.36	0.31	0.87	G5III-IV+A5+G5 III	5304	12	BOO 14	7+6	25 17	4+86	0.61	0+44	0+73	0+73	F8 IV
4534	BET	LEO 11 47+0	2 13	0.16	0.08	0.05	0.02	0.07	A3 V	5313	IDT	BOO 14	14+7	51 33	4+75	0+26	0+20	0+15	0+24	AT V
4540	BET	VIR 11 58+8	6 50	4.67	0.23	0.13	0.16	0.20	A4 V	5321	LAM	BOO 14	14+9	46 16	4+18	0.13	0+08	0+02	0+05	GP
4546	8	CEN 11 49+1	-66 57	(14.63)	*	1.29	0.94	1.61	K4 III	5326/9	IDT	LOD 14	16+8	-45 53	(13.56)	*	-0.18	-0.09	-0+24	B3 IV
4550 S	GAM	IB30 11 50+7	38 30	5.49	0.92	0.75	0.66	1+10	G8 Vp	5326	IDT	VIR 14	13+9	-49 15	4+09	0+54	0+51	0+49	0+77	F7 [II-IV]
4552 D	BET	HYA 11 50+6	-33 41	(14.25)	*	-0.10	0.02	-0.06	BB V	5340	A	BOO 14	14+3	55 35	2+06	2+52	1+24	1+24	K2 III	
4623	ALF	CRV 12 2+6	-26 30	4.00	0.30	0.23	0+32	0.49	F2 V	5367	PSI	CEN 14	18+1	-37 42	(4+04)	*	-0.02	0+01	-0+03	A0 IV
4630	EPS	CRV 12 8+1	-22 24	2.98	2.82	1.34	0.93	1.58	K3 III	5370 V	TAU 2	LUP 14								

## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

(SP.	B.S.	NAME	R.A. (1960)	DEC.	V	U-V	B-V	V-I	M SP.	B.S.	NAME	R.A. (1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.								
V	5616	PSI	BDO 15	27	6	4.55	2.57	1.23	0.93	-1.58	K2 III	6129	UPS	OPH 16	25.6	-8	1.7	4.63	0.22	0.16	0.13	0.21	A m				
V	5625		LUP 15	6.0	-42	43	5.00	-	-0.12	-0.07	-0.20	B7 Vinn	6132	ETA	DRA 16	23.4	61	36	2.74	1.61	0.91	0.61	1.07	G8 III			
V	5626	D	LAM	15	6.1	-45	8	(4.64)	T	-0.20	-0.08	-0.22	B3 V	6134	ALF	SCO 16	27.0	-23	21	0.89	3.17	1.83	1.56	2.79	M2 I		
V	5630		DEL	15	5.5	25	1	4.93	0.41	-0.43	-0.38	-0.60	F5 V	6141	22	SCD 16	27.7	-25	2	4.79	-0.86	-0.07	-0.06	-0.20	B2 V		
V	5634	49	BDO 15	15	5.5	25	1	4.93	0.41	-0.43	-0.38	-0.60	B9 V	6143	W	N	SCD 16	28.8	-34	37	(4.23)	*	-0.17	-0.04	-0.19	B2 IV	
V	5646/7	KAP	LUP 15	9.1	-48	35	(4.71)	*	-0.03	-0.03	-0.05	-0.05	-0.05	B9 V	6143	W	N	SCD 16	28.8	-34	37	(4.23)	*	-0.17	-0.04	-0.19	B2 IV
II-III	5651	8	LUP 15	10.1	-44	21	(4.81)	*	-0.14	-0.09	-0.25	B3 III	6146	V	g	HER 16	27.3	41	58	5.01	2.69	1.53	2.52	4.75	M6-III		
V	5652	IOT	LUP 15	9.9	-19	38	4.54	-0.46	-0.09	-0.04	-0.13	A P	6147	PHI	OPH 16	28.8	-18	32	4.27	1.63	0.92	0.65	1.09	G8 III			
V	5653	1	LUP 15	12.2	-31	22	(4.90)	*	0.38	0.37	0.68	F0 I	6148	BET	HER 16	28.5	23	35	2.77	1.63	0.94	0.64	1.11	G8 III			
V	5654	DEL	BDO 15	13.9	33	28	3.48	1.63	0.96	0.73	1.24	G8 III	6149	D	LAM	OPH 16	28.9	2	4	3.83	0.03	0.01	-0.00	-0.02	A1 V		
V	5655 D	MU	LUP 15	15.7	-47	44	(4.26)	*	-0.07	-0.00	-0.05	Ban	6153	OMG	OPH 16	29.8	-21	23	4.45	0.24	0.12	0.12	0.15	A P			
V	5656 S	BET	LUP 15	14.9	-9	14	2.61	-0.48	-0.11	-0.04	-0.13	B8 V	6159	29	HER 16	30.7	LI	34	4.85	3.33	1.48	1.19	2.05	K4 III			
V	5658 W	2	LUP 15	15.3	-30	1	4.33	2.17	1.09	0.11	1.37	K9	6161	15	DRA 16	28.0	68	51	5.01	-0.16	-0.05	0.04	0.01	B9 IV			
V	5659	DEL	LUP 15	18.7	-40	30	(3.21)	*	-0.12	-0.11	-0.33	B2 IV	6165	TAU	SCO 16	33.4	-28	8	2.82	-1.27	-0.25	-0.11	-0.36	B0 V			
V	5670	PHI	1	LUP 15	19.2	-36	7	(3.51)	*	0.54	1.19	2.06	K5 III	6166	H	SCO 16	33.7	-35	11	(4.15)	*	1.59	1.21	2.20	M2 I		
V	5670 D	EPS	LUP 15	20.0	-44	33	(3.36)	*	-0.17	-0.10	-0.26	B3 IV	6168	SIG	HER 16	32.8	42	31	4.20	-0.12	-0.02	0.03	0.02	B9 V			
V	5712	PHI	Z LUP	15	20.6	-36	33	(4.53)	*	-0.16	-0.03	-0.17	B5 V	6175	ZET	OPH 16	35.0	-10	29	2.57	-0.82	-0.02	0.10	0.06	09.5 V		
V	5713	MU	BDO 15	23.0	37	31	4.32	0.35	0.31	0.32	0.46	F0 IV	6172	D	ZET	HER 16	39.8	31	40	2.81	0.86	0.55	0.54	0.40	GO V		
V	5715	GAM	UNI 15	20.4	71	59	3.05	0.16	0.05	0.11	0.17	A3 [I-III]	6176	20	ETA	HER 16	41.5	39	30	1.53	0.92	0.67	1.15	1.07	J7 III-IV		
V	5719		DEL	15	24.7	-36	38	5.41	*	-0.16	-0.09	-0.26	B5 V	6237	W	DEL	OPH 16	44.5	56	51	4.04	0.33	0.17	0.15	0.36	F2 V	
V	5744	IOT	DRA 15	24.0	59	6	3.31	2.39	1.16	0.78	1.38	K2 III	6241	V	EPS	SCO 16	47.6	-34	13	(2.28)	*	1.17	0.85	1.46	K2 III-IV		
Vn	5747	BET	CRB 15	26.2	29	15	3.66	0.40	0.30	0.19	0.23	F0 p	6243	20	OPH	16	47.6	-10	43	4.66	0.54	0.48	0.44	0.69	F5 IV-V		
IV	5748	ZET	CRB 15	26.3	-16	44	5.67	*	-0.14	-0.09	-0.23	B2 Vnn	6247	V	MU	1	SCD 16	49.2	-38	3	3.02	*	-0.23	-0.12	-0.31	B1-5 V	
V	5750	D	BDO 15	30.3	41	1	5.02	0.18	0.07	0.15	0.19	Azn	6252	MU	2	SCD 16	49.7	-37	57	(3.56)	*	-0.22	-0.13	-0.35	B2 IV		
V	5775	D	GAM	LUP 15	32.5	-31	3	12.77	*	-0.20	-0.15	-0.37	B2 Vn	6254	52	HER	16	48.1	46	4	4.83	0.10	0.08	0.10	0.11	A2p	
V	5777	37	LUP 15	32.0	-9	56	4.60	1.88	1.02	0.75	1.28	K1 IVa	6271	W	ZET	SCO 16	51.8	-42	18	3.59	2.97	1.36	1.12	1.80	K5 III		
V	5778	THE	CRB 15	31.3	31	30	4.13	-0.68	-0.13	-0.05	-0.16	B7nm	6281	IOT	OPH 16	52.1	10	14	4.38	-0.40	-0.08	-0.09	-0.17	BB IV			
V	5780	W	LUP 15	32.2	-9	4	5.12	*	-0.11	-0.04	-0.16	B7 IV	6279	KAP	OPH 16	55.8	9	26	3.19	2.33	1.16	0.83	1.39	K2 III			
V	5781	D	d	LUP 15	33.1	-44	50	(4.54)	*	-0.16	-0.05	-0.36	B5 V	6315	19	DRA 16	55.8	65	12	4.91	0.45	0.47	0.44	0.71	F6 V		
V	5787	GAM	LUP 15	33.3	-14	39	3.91	1.77	1.03	0.73	1.28	G5 II-IV	6322	V	EPS	UMT 16	49.9	82	7	4.23	1.45	0.70	0.70	1.17	G5 II		
V	5787/2	DEL	SER 15	32.9	10	39	3.79	0.38	0.26	0.22	0.33	F0 IV	6324	EPS	HER 16	58.8	30	59	3.92	-0.12	-0.01	-0.01	-0.05	B9.5 V			
S	5793	V	ALF	CRB 15	33.0	26	51	2.27	-0.06	-0.02	0.07	0.05	A0 V	6334	*			17	21	-36	(4.86)	*	0.28	0.28	0.45	B1-4b	
Vpn	5794	UPS	LUP 15	34.4	-28	0	3.56	2.98	1.39	1.00	1.71	K5 III	6337	*			17	13	14	9	4.98	3.54	1.59	1.43	2.65	M3 III	
V	5797	DUO	LUP 15	35.5	-27	1	4.21	1.43	1.01	1.71	MO III	6355	*			17	13	14	9	4.98	3.54	1.59	1.43	2.65	A3 IV		
Vp/B9 V	5812	TAU	LUP 15	38.2	-29	39	3.65	-0.86	-0.17	-0.12	-0.32	B2+5 V	6378	D	ETA	OPH 17	8.0	-15	41	2.42	0.13	0.05	0.05	0.05	A2+5 V		
V	5820	PSI	LUP 15	37.2	-34	17	(4.61)	*	1.00	0.66	1.12	G5	6380	ETA	SCO 17	9.3	-43	11	(3.33)	*	0.42	0.36	0.56	F0 IVa			
V	5825	8	LUP 15	38.4	-44	32	(4.63)	*	0.44	0.39	0.55	F5 V	6395	ZET	DRA 17	8.7	65	46	3.17	-0.54	-0.11	-0.05	-0.17	BB III			
V	5830	KAP	LUP 15	39.6	-19	33	4.71	3.55	1.58	1.26	2.21	K3 III	64060V	ALF	HER 17	12.8	14	26	3.05	2.46	1.45	2.11	4.23	M5 II-III			
V	5839	PSI	2	LUP 15	40.1	-35	13	(4.74)	*	-0.14	-0.10	-0.28	B6 V	6410	DEL	HER 17	13.4	24	53	3.12	0.19	0.07	0.05	0.08	A3 IV		
IV	5842	D	IOT	SER 15	39.8	19	48	4.53	0.06	0.05	0.04	0.03	A1 V	6415	D	OPH 17	14.6	-24	72	2.26	1.15	0.81	1.40	2.23	K2 III		
IV	5849	D	GAM	CRB 15	41.0	28	25	3.85	-0.05	-0.01	0.01	-0.01	A0 IV	6418	P1	HER 17	13.7	36	51	3.16	3.06	1.96	1.68	1.68	K3 II		
IV	5854	S	ALF	SER 15	42.3	6	2.65	2.41	1.17	0.81	1.37	K2 III	6431	V	g	HER 17	15.8	33	9	4.83	-0.91	-0.15	-0.06	-0.26	B3 A9		
V	5859	BET	SER 15	42.3	6	3.44	3.00	0.07	0.06	0.05	A0 V	6436	69	HER 17	16.3	37	20	4.66	0.02	0.03	0.09	0.09	A2 V				
I-11	5862	BET	SER 15	44.4	-13	33	3.64	0.15	0.07	0.06	0.09	A2 V	6445	XI	OPH 17	18.6	-21	24	4.38	0.35	0.41	0.37	0.59	F2 V			
V	5866	LAM	SER 15	44.9	7	29	4.44	0.70	0.61	0.51	0.83	GO V	6446	N	SER 17	18.6	-12	49	4.31	0.06	0.03	0.07	0.07	A1 V			
V	5879	KAP	SER 15	46.9	16	41	5.89	*	-0.02	0.03	-0.01	B5 V	6454	THE	OPH 17	19.6	-24	58	3.27	-1.00	-0.22	-0.11	-0.33	B2 IV			
V	5881	MU	SER 15	47.5	-3	19	3.53	*	-0.02	0.01	-0.01	B6 V	6535	BET	OPH 17	29.5	52	50	2.78	1.63	1.00	0.58	1.16	G2 II			
V	5882	CHI	LUP 15	48.4	-23	52	5.29	*	-0.04	0.01	-0.04	B2 Vnn	6537	W	SIG	ARA 17	32.7	-66	29	(4.58)	*	0.00	-0.01	-0.06	A0 IV		
IV	5890	THE	LUP 15	51.5	-17	37	4.15	1.84	1.01	0.72	1.24	KO III-IV	6546	W	SCD 17	33.8	-38	37	(4.28)	*	1.07	0.75	1.30	g0			
V	5894	CHI	HER 15	51.3	42	34	4.53	0.58	0.48	0.48	0.80	F9 V	6553	THE	SCD 17	34.4	-42	58	1.16	0.41	0.35	0.55	F0 Ib				
V	5897	D	LUP 15	52.7	-19	16	5.89	*	-0.02	0.03	-0.01	B5 V	6554	NU	1	RA	17	31.4	55	13	4.89	0.29	0.27	0.26	0.37	A m	
V	5908	RHO	SCD 15	57.5	-38	18	(3.00)	*																			

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## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

B.S.	NAME	R.A.(1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.	B.S.	NAME	R.A.(1960)	DEC.	V	U-V	B-V	V-R	V-I	MK SP.	
6832 W	ETA	SGR 18 15.0	-36 46	[3.10]	*	1.54	1.56	2.88	M3 II	7615	ETA	CYG 19 54.8	-34 59	3.93	1.91	1.03	0.75	1.27	KO III	
6842		SGR 18 15.4	-27 4	4.63	3.45	1.66	1.22	2.14	I5K	7618	60	SGR 19 56.6	-26 18	4.82	1.43	0.89	0.65	1.12	IGS	
6859	DEL	SGR 18 15.4	-29 51	2.69	2.89	1.38	1.00	1.68	K2 III	7619 D	PSI	CYG 19 54.6	-52 20	4.91	0.19	0.12	0.11	1.12	K3 IV	
6866	74	OPH 18 18.9	-3 21	4.84	1.54	0.91	0.67	1.12	G8 III	7635	GAM	SCE 19 57.0	-19 23	3.47	3.51	1.57	1.20	2.12	K5+ III	
6868	106	HER 18 18.6	21 56	4.96	3.58	1.59	1.34	2.34	H0 III	7650	62	SGR 20 0	0.2	-27 49	4.59	3.48	1.45	1.87	3.53	M4 III
6869	ETA	SEN 18 19.2	-2 55	3.23	1.60	0.95	0.69	1.19	K0 III-IV	7653	15	VUL 19 59.5	-27 39	4.65	0.34	0.18	0.15	0.24	A m	
6872	KAP	Lyr 18 18.5	36 3	4.34	2.36	1.17	0.86	1.41	K2 III	7678		VUL 20 3.1	32 6	5.65	0.09	0.55	0.53	0.94	B1+5 Ia	
6879 W	EPS	SGR 18 21.5	-14 24	[1.84]	*	-0.02	-0.00	-0.01	A0 V	7685	RHO	ORA 20 2.6	67 45	4.50	2.81	1.31	0.93	1.58	K3 III	
6884	ZET	SCT 18 21.5	-8 58	4.67	1.69	0.95	0.70	1.17	K0 III	7708	28	CYG 20 7.9	36 43	4.92	-0.89	-0.12	0.02	-0.09	B3 V	
6895	109	HEI 18 22.0	21 45	3.86	2.34	1.17	0.85	1.45	K2 III	7710	THE	AQL 20 9.2	-0 56	3.21	-0.19	-0.07	-0.12	B9.5 III		
6896 D	ZI	SGR 18 23.0	-20 35	4.00	2.26	1.31	1.09	1.94	K2 II	7724	RHO	AQL 20 12.4	15 4	4.95	0.10	0.09	0.10	0.10	A2 V	
6897 W	ALP	TEP 18 24.0	-46 0	[3.50]	-0.19	-0.13	-0.14	S3 III	7730	30	CYG 20 12.4	46 42	4.82	0.24	0.10	0.19	0.24	A3 III		
6913	LAH	SGR 18 24.5	-25 27	2.82	1.94	1.04	0.76	1.32	K2 III	7735 V	31	CYG 20 12.4	46 37	3.80	1.70	1.28	0.97	1.73	K2 +B3V	
6918 V	59	SER 18 25.2	0 10	5.21	0.71	0.50	0.49	0.87	G0+A2	7736	29	CYG 20 13.0	36 41	4.99	0.12	0.12	0.21	0.27	A2 III	
6920 D	PHI	ORA 18 21.3	71 19	4.21	-0.43	-0.10	-0.05	-0.16	Adp	7739		VUL 20 13.0	25 28	4.77	-0.91	-0.18	-0.08	-0.27	B3 V	
6923 D	39	ORA 18 23.3	58 47	4.98	0.13	0.08	0.04	0.05	A1 V	7740	33	CYG 20 12.5	56 27	4.30	0.19	0.11	0.13	0.19	A3 IV-V	
6927	CHI	ORA 18 21.8	72 43	3.57	0.44	0.48	0.45	0.75	F7 V	7741	22	VUL 20 13.8	23 23	5.17	1.74	1.04	0.74	1.23	G2 II	
6930	GAM	SCT 18 26.9	-14 36	4.71	0.11	0.07	0.08	0.13	A3 V	7744	23	TEP 20 14.1	27 41	4.52	2.37	1.26	0.96	1.66	K3 III	
6945	42	DRA 18 25.9	65 32	4.81	2.32	1.19	0.85	1.18	K2 III	7747	ALF L	CAP 20 15.5	-12 38	4.26	1.89	1.08	0.79	1.32	G3 II	
6951 W	THE	ORA 18 30.6	-42 21	(4.63)	*	1.01	0.89	1.19	G5 III	7750 D	KAP	CAP 20 10.3	77 36	4.39	-0.14	-0.05	-0.01	-0.07	B9 III	
6973	ALF	SCT 18 33.0	-8 16	3.85	2.87	1.34	0.98	1.65	K3 II	7751 V	OMI 2	CYG 20 14.3	47 35	3.98	2.56	1.52	1.20	2.12	K3 Ib-II+	
6978	45	ORA 18 31.8	57 0	4.80	1.05	0.81	0.53	0.84	F7 I	7754	22	VUL 20 15.8	-12 40	3.58	1.64	0.95	0.69	1.16	G9 III	
7001	ALF	LYR 18 35.8	38 45	0.00	0.03	0.00	-0.04	-0.07	A0 V	7763 V	P	CYG 20 16.3	37 54	4.80	-0.16	0.41	0.54	0.80	S P	
7020 V	DEL	SCT 18 40.1	-9 6	4.71	0.91	0.39	0.30	0.48	F3 II-IV	7767 D		VUL 20 16.7	40 36	5.84	-0.68	0.10	0.14	0.15	OB	
7029 W		18 41.6	-35 41	(4.87)	*	-0.14	-0.14	-0.30	B3 V	7770	35	CYG 20 17.1	34 51	5.16	1.13	0.63	0.56	0.99	F5 Ib	
7039	PHI	SGR 18 43.2	-27 2	3.17	-0.44	-0.11	0.01	-0.10	B8 III	7773	NU	CAP 20 18.4	-12 53	4.76	-0.15	-0.04	-0.00	-0.07	B9 V	
7056	ZET	LYR 18 43.4	37 38	4.35	0.35	0.20	0.16	0.24	A m	7775	BET	CAP 20 18.8	-14 55	3.08	1.06	0.79	0.53	1.05	gKo+LATE B	
7057	110	HER 18 43.9	20 38	4.19	0.49	0.47	0.39	0.65	F6 V	7796	GAM	CYG 20 20.8	40 8	2.23	1.21	0.67	0.50	0.84	F8 Ib	
7063	BET	SCT 18 45.1	-4 48	4.21	1.94	1.09	0.79	1.36	G3 II	7804	39	CYG 20 22.3	32 4	4.45	2.87	1.34	1.01	1.68	K3 III	
7064		18 44.5	26 37	4.84	2.42	1.20	0.88	1.49	K3 III	7822 D	RHO	CAP 20 26.6	-17 57	4.80	0.42	0.38	0.34	0.54	F2 IV	
7065 V	R	SCT 18 45.3	-5 44	5.20	3.10	1.46	1.06	1.84	G0ia+K0ib+7834	7834	41	CYG 20 27.8	30 14	4.02	0.59	0.40	0.37	0.60	F5 II	
7069	111	HER 18 45.3	18 8	4.36	0.20	0.13	0.10	0.11	A3 V	7844	OMG 1	CYG 20 28.9	48 49	4.95	-0.72	0.09	0.02	-0.09	B2 V	
7106 V	BET	LYR 18 48.6	33 19	3.43	-0.57	-0.01	0.13	0.15	B p	7847	44	CYG 20 29.5	38 48	6.21	1.74	1.00	0.88	1.55	F5 Iab	
7116	NU	1 SGR 18 51.7	-22 48	4.81	2.70	1.42	1.00	1.69	cK2	7850	THE	CAP 20 28.9	62 52	4.21	0.33	0.19	0.18	0.23	A m	
7121	SIG	SGR 18 52.8	-26 21	2.07	-0.95	-0.22	-0.11	-0.31	B2 V	7852	EPS	DEL 20 31.3	11 10	4.04	-0.60	-0.12	-0.02	-0.13	B6 III	
7125	OMI	DRA 18 50.6	59 20	4.67	2.23	1.18	0.90	1.54	K0 II-III	7866	47	CYG 20 32.1	35 17	4.63	2.36	1.61	1.30	2.29	K2 Ib+B	
7133 D	113	HER 18 53.0	22 35	4.59	1.28	0.79	0.65	1.10	A5+G5 III	7871	ZET	DEL 20 33.4	16 33	4.69	0.22	0.11	0.15	0.20	A3 V	
7137		18 52.2	30 39	4.93	1.48	0.90	0.87	1.13	G8 II	7882 D	BET	DEL 20 35.7	14 28	3.63	0.52	0.44	0.40	0.64	F5 IV	
7139	DEL	2 LYR 18 53.1	36 50	4.30	3.32	1.67	1.79	3.42	M4 II	7884	71	AOL 20 36.3	-1 14	4.33	1.65	0.96	0.68	1.14	GB III	
7141/2	THE	SER 18 54.2	4 9	4.07	0.26	0.17	0.17	0.26	A5 V	7891	29	VUL 20 36.7	21 4	4.82	-0.09	-0.02	0.02	-0.02	B9.5 V	
7150	XI	2 SGR 18 55.3	-21 10	3.10	2.33	1.18	0.80	1.38	K1 II	7905	ALF	DEL 20 37.8	15 46	3.77	-0.25	-0.05	0.03	-0.01	B9 V	
7157 V	ZET	LYR 18 54.1	-43 54	4.00	3.00	1.59	2.03	3.95	IM5	7924	ALF	CYG 20 40.0	45 8	1.25	0.12	0.09	0.12	0.22	A2 Ia	
7176	EPS	AOL 18 57.8	15 1	4.01	2.12	1.07	0.76	1.28	K2 II	7928	DEL	DEL 20 41.6	15 56	4.44	0.42	0.32	0.27	0.44	ATp III	
7178	GAM	LYR 18 57.4	32 38	3.23	-0.11	-0.03	-0.04	-0.04	B9 III	7936	PSI	CAP 20 43.7	-25 25	4.13	0.47	0.44	0.36	0.56	F5 V	
7180	UPS	DRA 18 54.9	71 15	4.82	2.25	1.15	0.85	1.41	K1 II	7939	30	VUL 20 43.1	25 8	4.90	2.38	1.20	0.85	1.43	K2 II	
7193	12	AOL 18 59.5	-5 48	4.01	2.13	1.08	0.79	1.33	K1 II	7942	52	CYG 20 44.0	30 34	4.22	1.94	1.06	0.78	1.30	KD III	
7194 D	ZET	SGR 19 0.0	-29 56	2.60	0.17	0.08	0.04	0.05	A2n	7947	ETA	CEP 20 44.5	61 59	3.91	1.36	0.85	0.68	1.16	F8IV-V/K2IV	
7217	OMI	SGR 19 2.3	-21 48	3.77	1.87	1.01	0.64	1.18	F8 II	7949	EPS	CYG 20 44.6	33 49	2.65	1.91	1.03	0.72	1.29	KO III	
7234	TAU	SIG 19 4.4	-27 44	3.31	2.33	1.20	0.88	1.48	K1 II	7950	EPS	AOR 20 45.5	9 39	3.77	0.01	-0.00	0.07	0.07	A1 V	
7235	ZET	SGR 19 5.3	-25 20	4.82	0.90	0.56	0.45	0.81	F5	7953	OMG	CAP 20 49.4	-9 11	4.43	1.81	1.47	1.27	2.19	M1 III	
7236	LAM	AOL 19 4.1	-6 57	3.43	-0.37	-0.11	-0.03	-0.11	B9 V	7955		20 44.4	57 26	4.52	0.64	0.54	0.47	0.75	F8 IV-V	
7254 W	ALF	CRA 19 6.7	-37 58	(4.10)	*	0.03	0.04	0.04	A2n	7957	ETA	CEP 20 44.8	15 59	3.91	1.36	0.85	0.68	1.16	KO IV	
7259	BET	CRA 19 7.2	-39 24	(4.10)	*	1.18	0.82	1.43	G5	7963 D	LAM	CYG 20 45.8	32 21	4.54	-0.61	-0.12	-0.03	-0.14	B5 V	
7264	PI	CRA 19 7.4	-21 5	2.87	0.57	0.33	0.34	0.59	F2 II-III	7977	55	CYG 20 47.5	48 58	4.87	-0.02	0.43	0.45	0.76	B3 IIa	
7292 D	PSI	CRA 19 13.1	-25 20	4.82	0.90	0.56	0.45	0.81	F5	7983	OMG	CAP 20 49.4	-27 4	4.13	3.47	1.62	1.25	2.19	M1 III	
7298	ETA	LYR 19 12.4	39 5	4.38	-0.81	-0.14	-0.11	-0.26	B2 IV	7990	MU	AOR 20 50.5	-9 8	4.73	0.42	0.32	0.26	0.41	ASm	
7306	1	VUL 19 14.5	21 20	4.77	-0.59	-0.04	0.03	-0.05	B3 IV	7994	31	VUL 20 50.4	26 56	4.61	1.28	0.82	0.68	1.14	GB III	
7310	DEL	DRA 19 1																		

## THE ARIZONA-TONANTZINTLA CATALOGUE—MULTICOLOR PHOTOMETRY UBVRI

SP.	B.S.	NAME	R.A. (1960) DEC.	V	U-V	B-V	V-R	V-I	MK SP.	B.S.	NAME	R.A. (1960) DEC.	V	U-V	B-V	V-R	V-I	MK SP.		
III	5116 V	MU	CEP 21 42.2	38 36	4.13	4.59	2.26	2.07	3.81	K2 Ia	8720 W	DEL PSA 22 53.7	-32 45	14.20	*	0.97	0.74	1.35	G4	
	5117 II	CEP	21 43.3	38 11	4.57	2.19	1.11	0.84	1.38	K0 III	8726	PSA 22 54.7	49 31	14.94	3.72	1.78	1.34	2.39	K5 Ib	
IV	5322 V	DEL	CAP 21 44.8	-16 19	2.81	0.40	0.29	0.24	0.40	A6m	8728	ALF PSA 22 55.4	-29 50	1.15	0.14	0.08	0.11	0.13	A3 V	
III	5327 NU	CEP	21 43.8	62 17	5.92	-0.33	0.31	0.28	0.45	O9 II	8729	PEG 22 55.4	20 33	5.45	0.65	0.66	0.54	0.48	DG0	
III	5334	NU	CEP 21 44.3	60 56	4.29	0.64	0.51	0.31	0.94	A2 Ia	8748 W	PSA 22 54.8	84	4.72	3.13	1.43	1.08	1.81	K4 III	
I	8135	PI	Z CYG 21 45.3	49 7	4.24	-0.84	-0.12	-0.04	-0.17	B3 III	8752	PSA 22 58.4	36 43	5.13	2.88	1.55	1.18	2.03	G0 Ia	
IV	4103 W	GRU	21 51.5	-37 33	(3.00)	*	-0.10	-0.05	-0.11	B5 III	8752	PSA 22 58.4	42 7	3.62	-0.62	-0.09	0.01	-0.07	B6p	
III	5371 L3	CEP	21 53.5	56 25	5.81	0.68	0.71	0.63	1.28	K0 IIb	8773	BET PSC 23 1.8	3	3.36	4.52	-0.62	-0.12	-0.02	-0.15	B5pe
V	5383 V	VV	CEP 21 55.5	63 26	4.98	2.13	1.75	1.70	3.04	H2+Ia-lab	8775 V	BET PEG 23 1.8	27 52	2.42	1.62	1.67	1.51	2.83	M2 II-III	
IV	5452 OMI	AOR	22 1.2	-22 21	4.70	-0.48	-0.07	0.00	-0.05	B6 V	8780	AND 23 2.4	49 49	4.66	1.94	1.07	0.74	1.31	K0 III	
V	8411 W	LAM	GRU 22 3.7	-39 44	(14.44)	*	1.37	1.00	1.78	MO III	8781	ALF PEG 23 2.8	14 59	2.47	-0.06	-0.03	0.01	-0.01	89+5 III	
III	8413 NU	PEG	22 3.6	4 51	4.03	3.25	1.46	1.07	1.83	K4 III	8789	B6 AOR 23 4.5	-23 58	4.48	1.49	0.90	0.71	1.20	I9	
III	+By+	8414 ALF	AOR 22 3.7	-0 31	2.92	1.77	0.98	0.66	1.13	G2 IIb	8795	PSC 23 5.0	9 11	4.51	3.53	1.59	1.26	2.28	M2 III	
III	8417 D	XI	CEP 22 2.6	64 25	4.29	0.43	0.34	0.32	0.46	A m	8796	PEG 23 5.1	25 15	4.77	2.50	1.35	0.97	1.65	K0 Ibp	
V	8418 IOT	AOR	22 4.3	-14 4	4.25	-0.34	-0.07	0.00	-0.13	B8 V	8797	AND 23 4.9	59 12	4.88	-0.89	-0.02	0.10	-0.01	80.5 IV	
IV+V	8425 W	ALF	GRU 22 5.7	-47 9	(11.73)	*	-0.17	-0.08	-0.14	B5 V	8808W	PSA 23 6.1	63 26	5.25	-0.61	-0.01	0.05	-0.01	B3 III	
IB	8428 L9	CEP	22 3.9	62 5	5.10	-0.76	0.07	0.14	0.19	99.5 1b	8812	B6 AOR 23 7.3	-21 23	3.64	2.45	1.23	0.84	1.44	K0 III	
III	8430 IOT	PEG	22 5.1	25 9	3.76	0.45	0.45	0.39	0.64	F5 V	8817	B9 AOR 23 7.4	-22 50	4.68	1.02	0.45	0.55	0.91	SG2 III	
IB	8431 W	MU	PSA 22 6.1	-33 11	(14.49)	*	0.07	0.07	0.12	A2 V	8819 D	PI CEP 23 6.6	75 11	4.42	1.27	0.82	0.65	1.08	G2 III	
III	8450 THE	PEG	22 8.1	5 59	3.55	0.17	0.08	0.05	-0.09	A2 V	8820 IOT	GRU 23 8.1	-5 28	(3.88)	*	1.01	0.75	1.31	K0 III	
IB-II+A	8454 PI	PEG	22 8.2	32 59	4.27	0.65	0.48	0.39	0.66	F5 II-III	8830	7 AND 23 10.7	49 12	4.53	0.34	0.30	0.30	0.47	F0 IV	
III	8465 ZET	CEP	22 9.3	58 1	3.16	3.27	1.55	1.08	1.86	K1 Ib	8832 S	23 11.3	36 57	5.57	1.90	1.01	0.83	1.33	K3 V	
J	8468 24	PEG	22 9.0	2 9	4.49	1.53	0.92	0.49	1.17	G8 III	8834	PHI ADR 23 12.3	-6 16	4.23	3.45	1.55	1.28	2.37	M2 III	
IB	8469 LAM	CEP	22 10.1	59 13	5.04	-0.51	0.23	0.26	0.41	O6I	8841 PSI 1 ADR 23 13.4	-9 18	4.25	2.12	1.11	0.79	1.35	K0 III		
IB	8485	PEG	22 12.2	39 31	4.49	2.82	1.38	1.01	1.74	K3 III	8852 GAM	PSC 23 15.1	3 4	3.69	1.46	0.91	0.71	1.23	GT III	
V	8594 EPS	CEP	22 13.4	56 51	4.20	0.33	0.27	0.29	0.44	F0 IV	8858	PSI 2 AOR 23 15.8	-9 24	4.40	-0.69	-0.14	-0.05	-0.19	B5 V	
+LATE B	8593 1	LAC	22 14.2	37 33	4.15	3.06	1.46	1.04	1.75	K3 II-IV	8860	8 AND 23 15.9	48 48	4.86	3.64	1.67	1.46	2.72	H2 II	
IB	8599 THE	AOR	22 14.7	-7 59	4.15	1.80	0.99	0.73	1.20	G8 II-IV	8863	GAM SCL 23 16.7	-32 45	(4.04)	1.40	1.15	0.84	1.47	GG III	
III	8618 GAM	AOR	22 19.6	-1 35	3.85	-0.18	-0.05	0.04	0.00	89 III	8872 D	OMI CEP 23 16.9	75 54	4.76	1.33	0.84	0.65	1.10	K0 III	
IV	8620 31	PEG	22 19.5	12 0	5.02	-0.92	-0.15	-0.04	-0.20	82 V	8880 TAU	PEG 23 18.7	23 31	4.60	0.30	0.17	0.20	0.29	A5 IV	
II	8522 32	PEG	22 19.5	28 8	4.81	-0.19	0.00	0.05	-0.08	8811	8892	AOR 23 20.9	19 19	3.97	2.06	1.10	0.82	1.42	K0 III	
IIab	8523 2	LAC	22 19.4	40 20	4.36	-0.50	-0.10	-0.01	-0.12	8814	8905 UPS	PEG 23 23.4	23 11	4.44	0.79	0.61	0.56	0.86	F8 IV	
III	8538 BET	LAC	22 22.0	52 2	4.44	1.79	1.03	0.76	1.32	G9 III	8906	99 AOR 23 24.4	-20 51	4.38	3.29	1.49	1.12	1.93	K5 III	
III	8539 PI	AOR	22 23.2	1 10	4.64	-0.79	-0.09	0.15	0.15	81 nekk	8911 KAP	PSC 23 24.9	1 2	4.94	0.00	0.04	0.00	0.04	A2p	
III	8541 4	LAC	22 22.9	49 16	4.58	0.25	-0.09	0.14	0.24	8916 THE	PEG 23 25.9	6 10	4.30	2.11	1.08	0.80	1.33	K1 III		
IB+B	8551 35	PEG	22 25.8	4 31	4.80	1.93	1.05	0.77	1.31	K0 III-IV	8923	70 PEG 23 27.1	12 32	4.56	1.69	0.94	0.74	1.19	GB III	
V	8556 W	DEL 1	GRU 22 26.9	-43 42	(3.96)	*	1.02	0.73	1.34	G2 II	8926 V	AR CAS 23 28.5	58 20	4.54	-0.76	-0.12	0.03	-0.12	83 V	
IV	8558/Z	ZET	AOR 22 25.8	-0 13	3.66	0.37	0.40	0.37	0.61	F2 IV	8937	BET SCL 23 30.8	-38 2	(4.36)	-0.09	-0.00	-0.07	0.07	A p	
IV	8558/W	DEL 2	GRU 22 27.4	-43 57	(4.10)	*	1.56	1.73	3.32	M4	8939	101 ADR 23 31.2	-21 8	4.72	0.03	0.04	0.09	0.12	Adm	
V	8571 V	DEL	PEG 22 27.7	58 13	4.34	1.52	0.88	0.74	1.19	G1 Ia	8949	PEG 23 32.9	-42 50	(4.70)	*	0.11	0.12	0.19	A2 V SF	
V	8572 D	5	LAC 22 27.9	47 30	4.37	2.78	1.68	1.39	2.46	MO Ia	8950	PSA 23 35.7	-45 43	14.73	*	0.09	0.11	0.16	A2 V	
II	8573 SIG	AOR	22 28.5	-10 53	4.82	-0.18	-0.06	-0.01	-0.04	AO IV	8961 V	LAM AND 23 35.6	46 14	1.74	1.01	0.78	1.35	GB III-IV		
III	8576W BET	PSA	22 29.2	-32 33	(4.28)	*	0.01	0.02	0.04	AO V	8965 IOT	AND 23 36.2	43 3	4.29	-0.39	-0.10	0.02	-0.07	B8 V	
V	8579 6	LAC	22 28.7	42 55	4.52	-0.85	-0.11	0.00	-0.11	B2 V	8969	IOT PSC 23 37.9	5 25	4.13	0.53	0.51	0.43	0.72	F7 V	
V	8582 EPS	PSA	22 38.4	-27 15	4.16	-0.44	-0.12	-0.05	-0.15	B8 V	8988 OMGS 2 AOR 23 40.6	-14 46	4.51	-0.17	-0.04	0.03	-0.02	BB.5 V		
III	8632 11	LAC	22 38.7	44 4	4.46	2.70	1.33	0.93	1.61	K3 III	8997 D	PEG 23 42.0	78 8	4.94	1.60	0.95	0.71	1.19	K1 III	
IV-V	8534 ZET	PEG	22 39.7	10 37	3.19	-0.35	-0.10	-0.03	-0.10	BB V	8997 KAP	AND 23 38.4	44 7	4.15	-0.30	-0.08	0.02	-0.05	B8 V	
IV	8562 9	LAC	22 40.6	51 21	4.63	0.35	0.24	0.24	0.38	A7 IV	8998 104 ADR 23 39.7	-18 2	4.84	1.31	0.81	0.62	1.04	G0 Ib		
IV	8564 OM1	PEG	22 39.9	29 6	4.79	-0.01	-0.01	0.05	0.04	A1 V	8998 LAM	PSD 23 40.0	1 34	4.49	0.32	0.22	0.19	0.29	A7 V	
IV	8564 W	RHO	GRU 22 41.2	-41 37	(14.84)	*	1.01	0.76	1.41	G3	9066 V	R CAS 23 56.3	51 10	6.75	1.71	1.00	0.72	1.25	M7e	
IV	8565 ETA	PEG	22 41.1	30	5 59	1.94	1.46	0.87	1.11	G4 II+FO	9071 D SIG	CAS 23 56.9	55 32	4.29	-0.90	-0.08	0.05	-0.05	B1 V	
III	8665 XI	PEG	22 44.7	11 59	4.19	0.47	0.51	0.44	0.75	F7 V	9072 OMGS	PSC 23 57.3	6 39	4.01	0.49	0.41	0.38	0.62	F4 IV	
V	8667 LAM	PEG	22 44.6	23 21	3.95	1.96	1.07	0.75	1.27	G3 II-III	9073 30 PSC 23 59.9	-6 14	4.41	3.47	1.63	1.56	2.97	M3 III		
V	8679 TAU	AOR	22 47.5	-13 48	3.98	3.53	1.59	1.18	2.13	MO III	9091 ZET	SCL 23 59.7	-29 57	4.78	-0.70	-0.15	-0.05	-0.19	B5 IV	
V	8684 MU	PEG	22 48.1	24 23	3.48	1.81	0.94	0.68	1.15	G3 II	9098 2 CET	PSL 23 60.7	-17 34	4.56	-0.16	-0.04	0.03	0.00	B9.5 V	
V	8684 IOT	PEG	22 48.2	65 59	3.53	1.96	1.05	0.83	1.34	K3 III	9099 7	PSL 23 60.7	-17 34	4.56	-0.16	-0.04	0.03	0.00	B1 V	
IV:e	8695W GAM	PSA	22 50.3	-33 5	(4.46)	-0.05	0.02	0.04	AO V	9102	OMGS	PSC 23 57.3	6 39	4.01	0.49	0.41	0.38	0.62	F4 IV	
V	8698 LAM	AOR	22 50.5	-7 48	3.79															

1983 Mar. 13 25th MSS

## 流星の写真測光 IV.

小笠原 雅弘

写真測光用の比較星としては Arizona-Tonantzintla Catalog (ATCと略す) が用いられる。5mag. までのデータは ATC に、暗い方は SAO の等級 Mu を用いてはならないことがある。SAO の Mu 等級はさほどすぐ source-catalog から寄せ集めであり、精度がよくないとわれている。そこで ATC と SAO を比べることにする。SAO Mu 等級の精度を見積る。

(4) B. Iriarte, H.L. Johnson, R.J. Mitchell and W.K. Wisniewski, 1965, Five color photometry of bright stars. Sky Telesc., July, p.21.

は便利で楽しい表を載せている。全体で11頁、表の部分は7頁たるべで、1234個の星の観測光値が並べてある。 $-50^{\circ}$  より北の5.0等以上の星を網羅しているが、変光星や特殊な星でもっと暗いものがいくつかまざっている。 $\chi$  Cyg 10.31, Barnard's star 9.54,  $\alpha$  Cet 8.71など。もともと Johnson が昔この雑誌に50個の明るい星の表を出したものの拡張なので、明るさの点で物足りないが、統一されたシステムでこれだけの数が集められており、手に入りやすいという特徴がある。内容は星の番号(BS=HR), 星の名前 (Beyer または Flamsteed), 1960年の座標,  $V$ ,  $U-V$ ,  $B-V$ ,  $V-R$ ,  $V-I$  と、スペクトルの MK 分類型である。 $V$ を基準にして他を表わすということで  $U-B$  があらわに出ていないのが少し不便かもしれない。観測に使うときは座標が出てるのは便利だし、S20の光電管の範囲はこの表だけで済むのもうまい。これだけ連続光の情報が与えられており、スペクトル型で線スペクトルのことわざがあるので、小天文学の楽しみが味わえる。カタログを見る楽しみの一つに注記がある。BS番号と星名の間に、標準星とか変光星とか重星とかの印があり、表の前後に簡単なノートがある。明るい星の観測プログラムを作るときなどはこの表だけで結構楽しめるのである。

of the Bright Stars. にまとめられている。

### —SAO データ—

SAO には ATC に含まれる星 129 個を全天より選び Mu 等級を比べる。SAO では Mu 等級の source catalog は 10 を越えず  $3.0^m$  ~  $5.0^m$  までの 27 の source によっている。

### — ATC —

星団星表めぐり 16. 星の明るさと  
色の星表 近藤 雅良. 手

Catalina Observatory (Arizona-  
Univ.)

28 inch - カセグイン鏡  
Mexico National Observatory  
(Tonantzintla)

40 inch - カセグイン鏡  
の 2 台を用いて 1963 - 1965 年  
にかけて測定されており  
各個の測光精度は Catalina  
 $\pm 0.017 \text{ mag} (\text{Mu})$ , Tonantzintla  
 $\pm 0.035 \text{ mag} (\text{Mu})$  と発表  
されている。

SAO 観測のより詳しい結果は  
H.L. Johnson et al. (1966)  
UBVRIJKL Photometry

カナダコトハ  
H: Henry Draper Catalogue

A.J. Cannon and E.C. Pickering (1918-24)

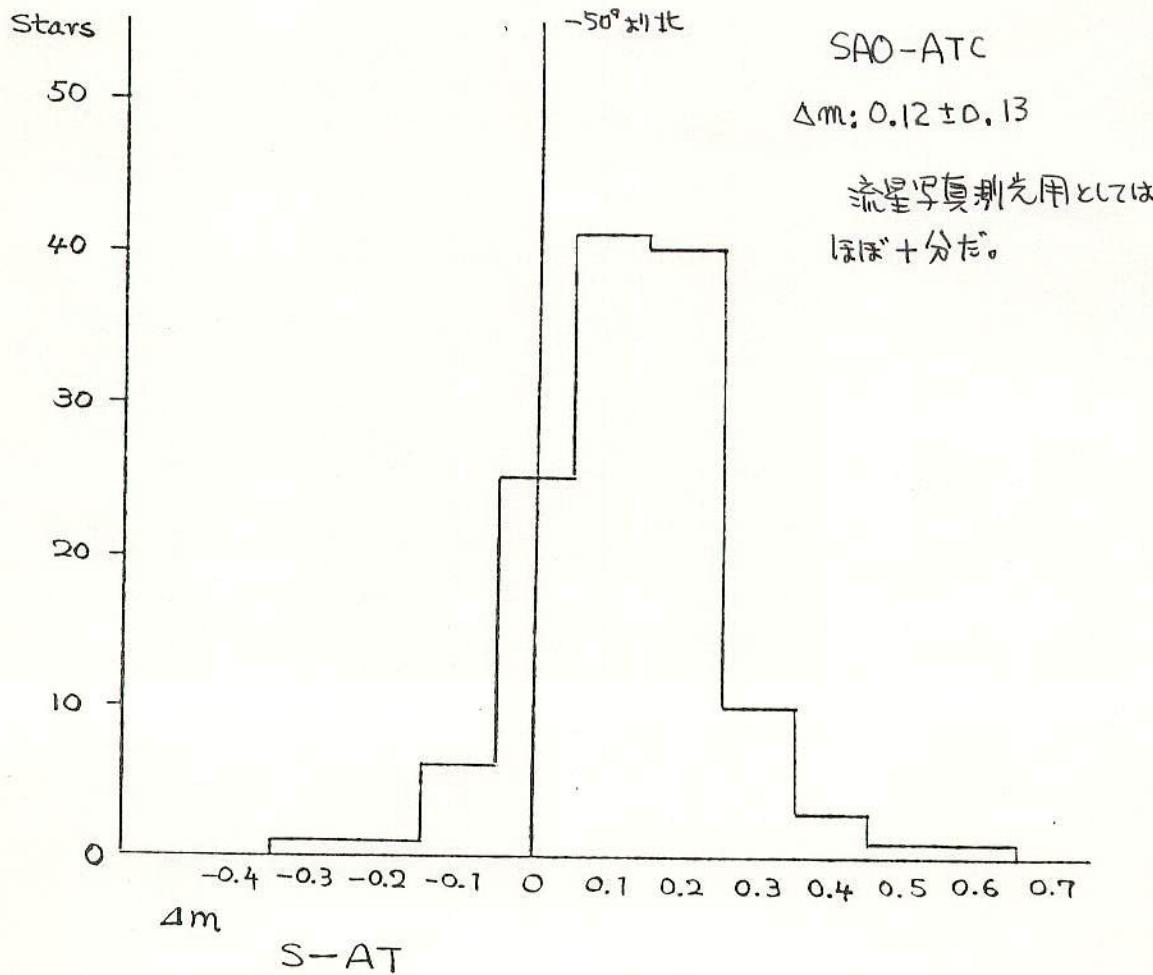
Ann. Astron. Obs. Harvard College

T: Harvard Photometry or measured at San Luis  
Argentina (G.C.)

Fig. 1

## Magnitude Difference

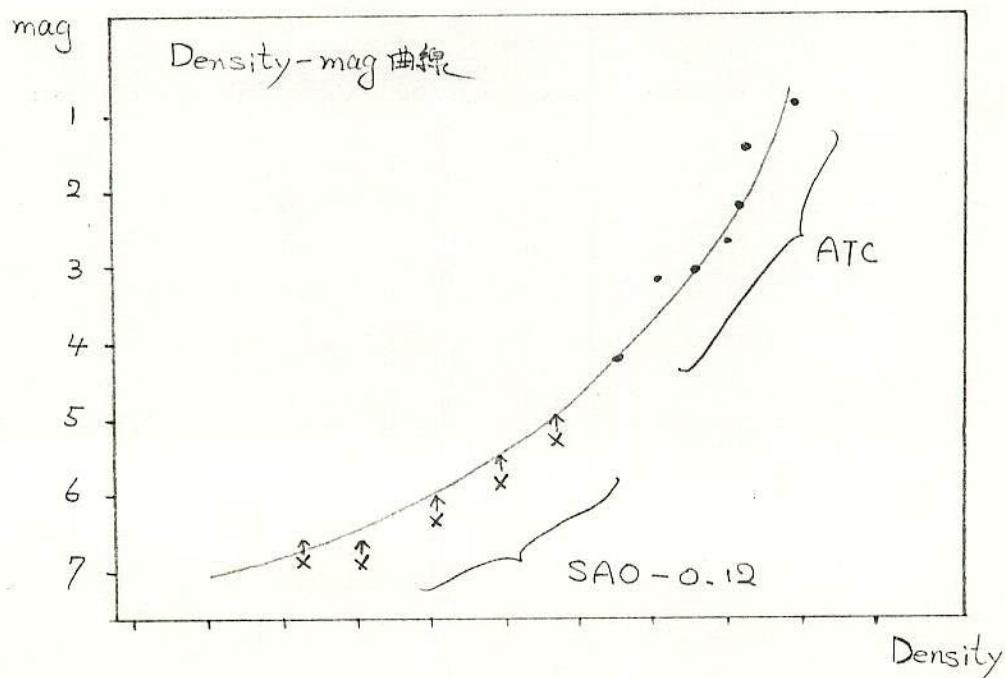
Arizona Tonantzintla — SAO

 $\sim 5 \text{ mag}$   
1200 $\pm$  $\sim 9 \text{ mag}$   
265 $\pm$ 

SAO(主に H.T) は ATC より平均 0.12 mag. 暗いほど  
かどる。ATC を正しいものと考へると SAO の等級誤差は (S.D.)  
 $\pm 0.13 \text{ mag}$  と考えられる。

したがって、測光用比較星を ATC と SAO どちら採る場合  
SAO - 0.12 とし、ATC は準定して等級に直して用いるべきである。

なお、これでも SAO では約 10% の星が写真測光の 2 つの  
尺度である。0.2 mag 以上の誤差を含むので流星等の等級を  
決定するには数多く (100 個以上) の恒星を用いる方がよい。



次回の Icarus Comet Special issue の Index

## Volume 47, Number 3, September 1981

Special *Icarus* "Comet" Issue

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The Subject Index for Volume 47 will appear in the December 1981 issue as part of a cumulative index for the year 1981.

Omar — M. Ogasahara

## 流星痕の色とカラー・フィルムの再現性.

田口泰雄 (信州大OB)

1983.03.13

流星物理セミナー

- サクラカラーフィルム 400 (ニュータイプ)
- フジクローム 400
- エクタクローム 400
- コダカラーフィルム 400
- コダカラーフィルム II

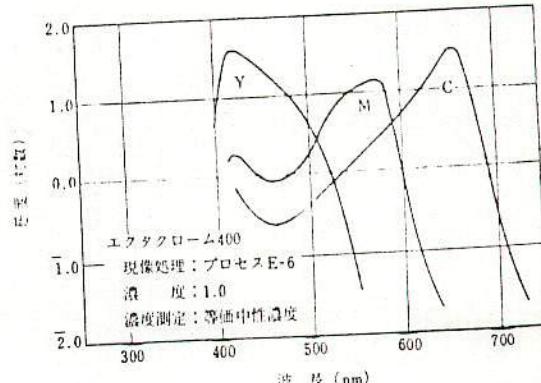
## の分光感度曲線

1982 Ori痕は、写真ではオレンジに  
写っているが、眼視では青とオレンジが  
混った色に見えた。

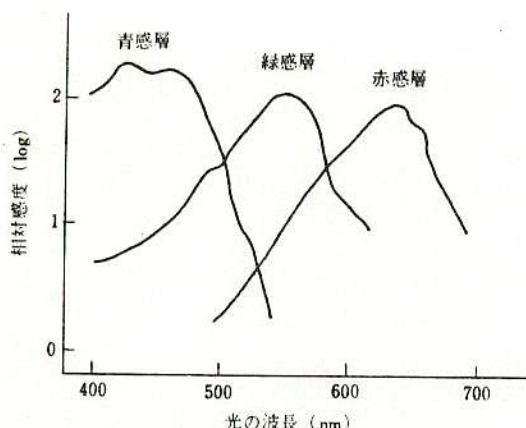
写真の色は赤オレンジ、眼視は青にオレンジ  
又は緑にオレンジが混った色に見えた。なぜ  
遠くで見えたのかを考えている。

日没10分後の地平線	1/60秒	F5.6
スポットライト照明のサークル	1/250秒	F2.8
夜の室内	1/30秒	F2.8
ステージショー	1/60秒	F2.8
夜景	1/60秒	F2.8
夜のネオン	1/125秒	F4
ナイター試合	1/125秒	F2.8

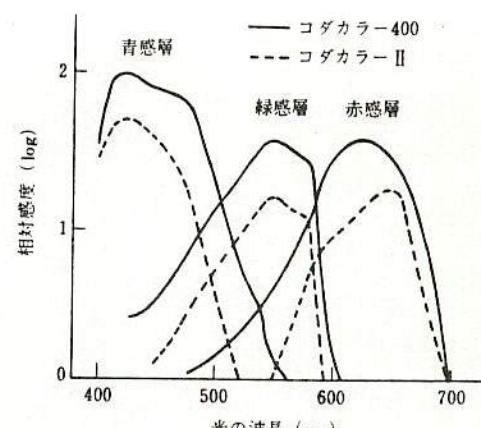
現像処理 プロセスE-6、キットにより自家現像処理が可能。



エクタクローム400の分光感度曲線

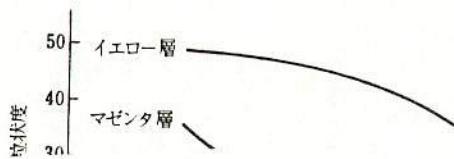


(a) サクラカラー400新タイプの分光感度



(b) コダカラーフィルムの分光感度

図4 サクラカラー400新タイプとコダカラーフィルムの分光感度



タクローム400も、高感度にもかかわらずナチュラルな色再現性は高く評価されよう。フジクローム400はカブリ濃度もエクタクローム400より少なく、スケのよいすっきりした画面となるが、最暗部の黒は緑味を帯びる。

ASA 400という高感度フィルムに対してストロボ照明の機会は少ないが、フジクローム400の場合、肌色再現に関しては、ストロボ光源との合性はよくないようだ。しかし、螢光灯、ミックス光に対しては、他のフィルムより合性はよいといえよう。メーカー推薦の螢光灯に対する補正是表2、また相反則不軌の補正是表1のごとく示されている。

### ●特性曲線

フジクローム400のE-6処理による特性曲線は図4のごとくで(図5はエクタクローム400)、良好なカラーバランスを得ている。図6は今回検討した3種類のフィルムの視覚黒濃度による比較で、フジクローム400とエクタクローム400の階調再現性はほとんど同一で、感度も等しい。得られる最大黒濃度は三者ともほぼ3.0で、カブリ濃度はエクタクローム400がわずかに高い。

たように、分光感度分布をある程度コダカラーフィルムのように中央によせた形とし、異種光源特性を向上させ、タイプGに近似したタイプとしている。しかし、データライドタイププレミアムは、

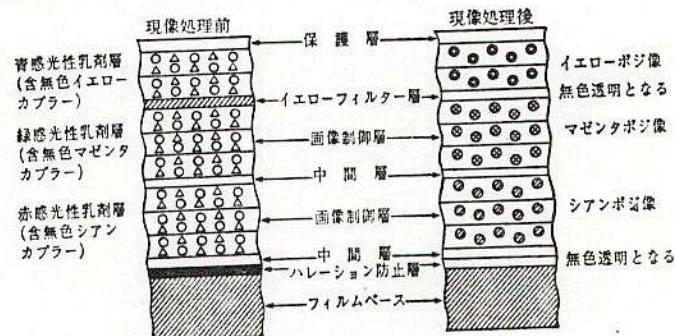


図1 フジクローム400の層構成

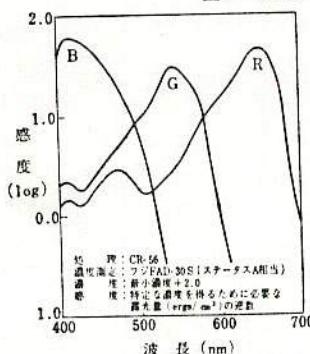


図2 分光感度曲線

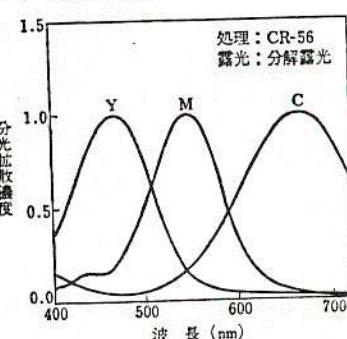


図3 色素の分光感度

表1 相反則不軌の露出補正

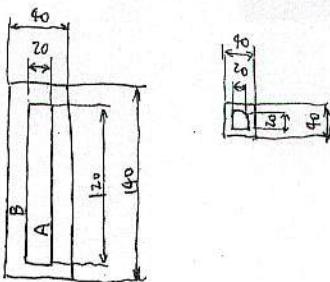
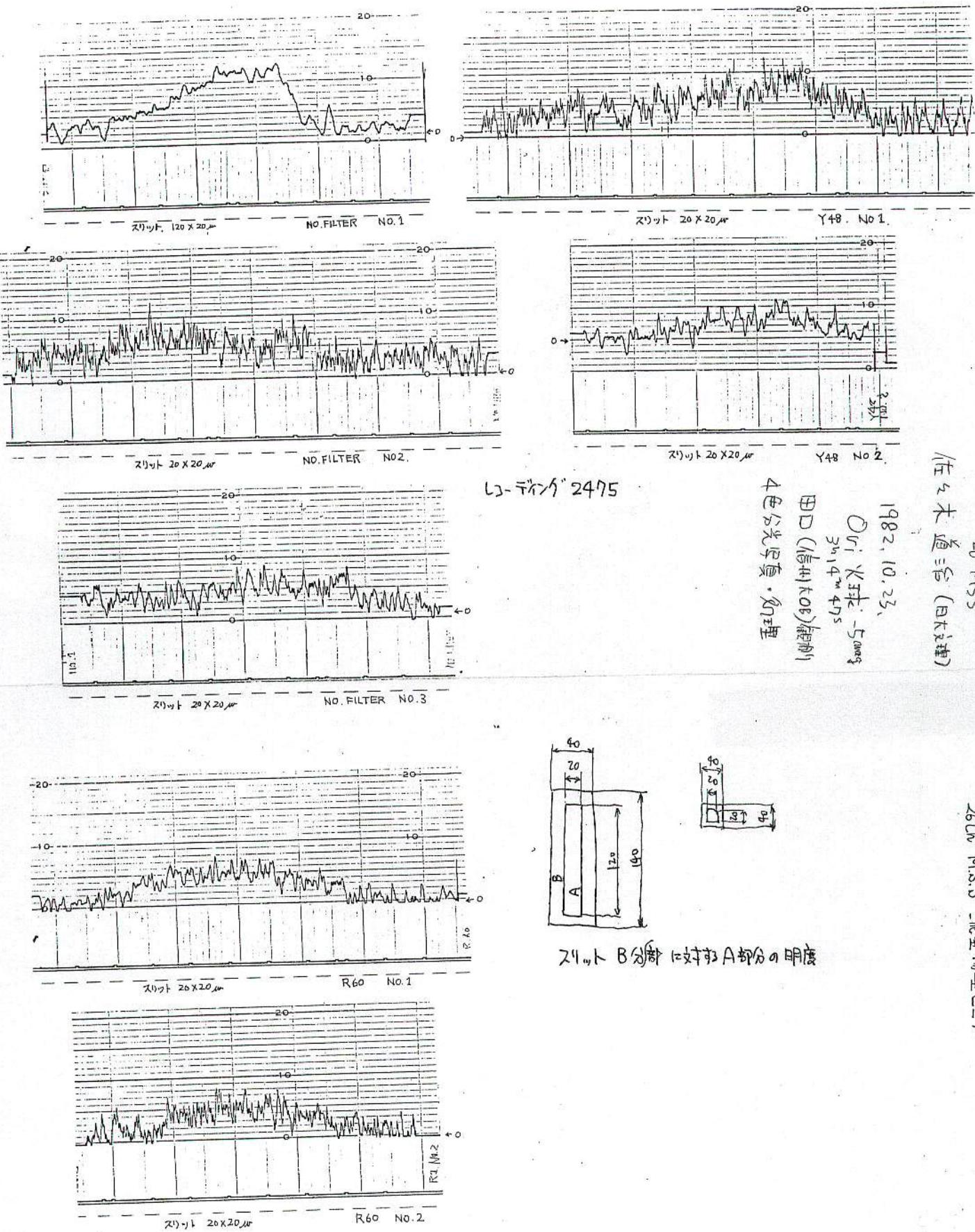
露出時間(秒)	1/1000~1	4	16	32
色補正フィルター	不要	5°C	10°C	15°C

26 MSS

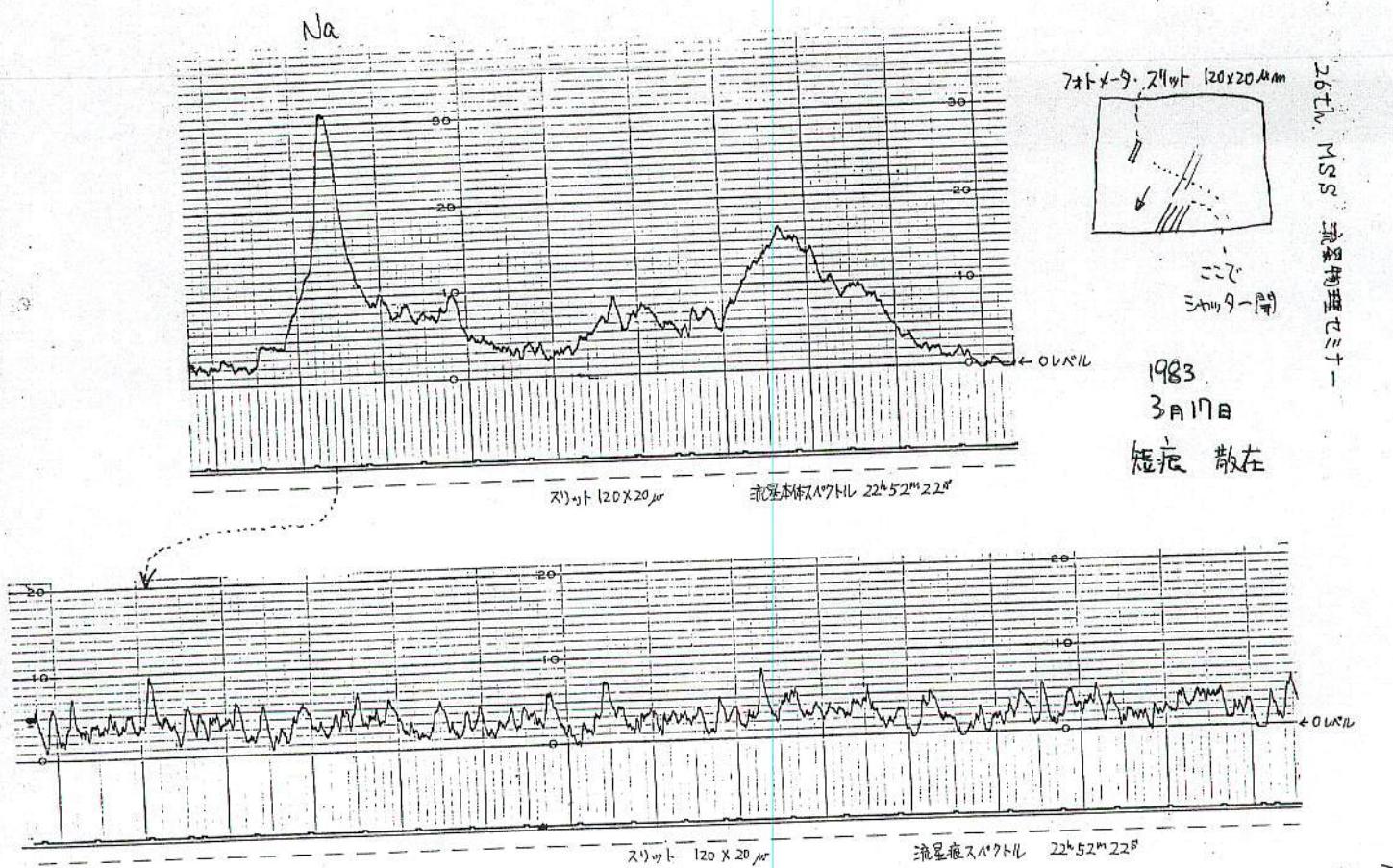
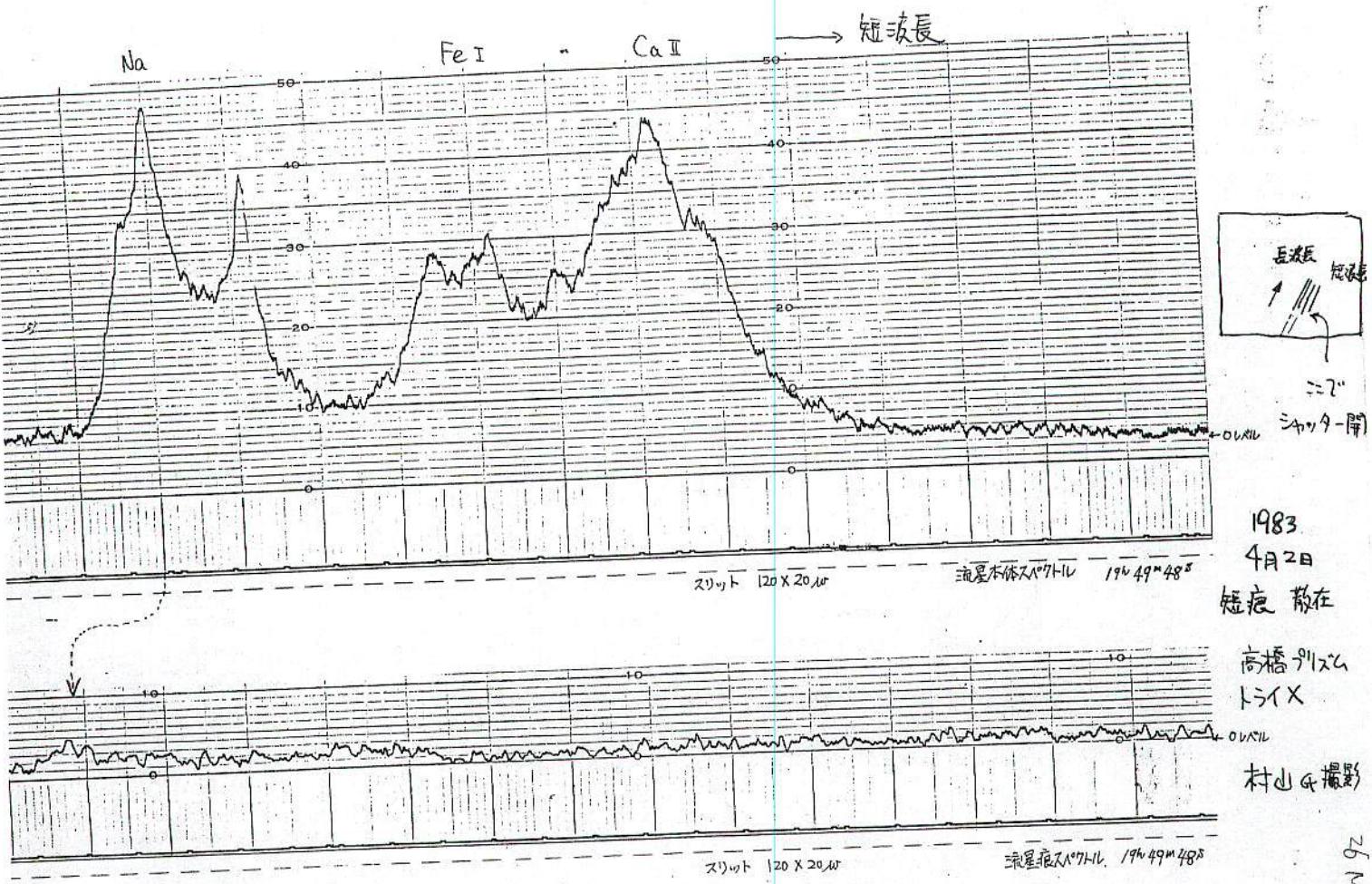
佐々木道三(日大物理)

26th M.S.S. 流星物理セミナー

MSS-026



24. ト B 分割 = 支持 A 部分の明度



MSS-026